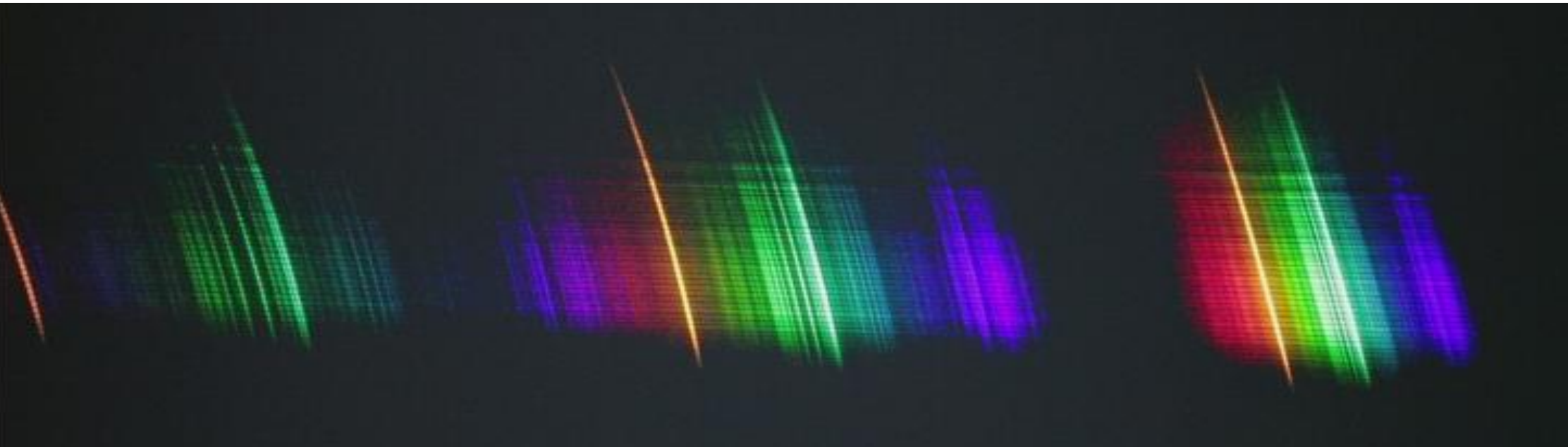


Meteor Spectroscopy, OHP 2018

Martin Dubs, images by Koji Maeda

SAG, FMA



Content

- Meteor astronomy
 - Observation
 - Network
 - Video observation: hardware and software
- Meteor spectroscopy
 - Hard- Software
 - Wavelength calibration
 - Spectrum extraction
 - Instrument response
 - Summary

Origin

- Comets
 - Dust tail crossing earth orbit
 - Dusty snowball
 - Low density
 - Meteor streams
 - 80 – 120 km height
- Asteroids
 - Stone or iron meteorites, higher density, material strength
 - Reaches lower atmosphere
- Very old material from beginning of solar system, original state
- Space experiment for the amateur

Observation

- Visual
 - Number, brightness, direction
- Photographic
 - Radiant with simultaneous observation
- Video
 - Path, velocity with simultaneous observation
 - Origin, extrapolate backwards: Orbit elements
 - Meteorites: dark flight, extrapolate forward
 - drop or impact region

<http://www.meteorastronomie.ch>

■ Swiss meteor network

Fachgruppe Meteorastronomie

HOME / NEWS

FACHGRUPPE

STATIONEN DER FMA

WAS SIND METEORE?

FOTO / VIDEO

RADIO

SPEKTROSKOPIE

VISUELL

MOND-METEORE

HISTORISCHES

TLE

AGENDA

ERGEBNISSE

BILDER

UNTERLAGEN

LINKS

KONTAKT

MITGLIEDERBEREICH

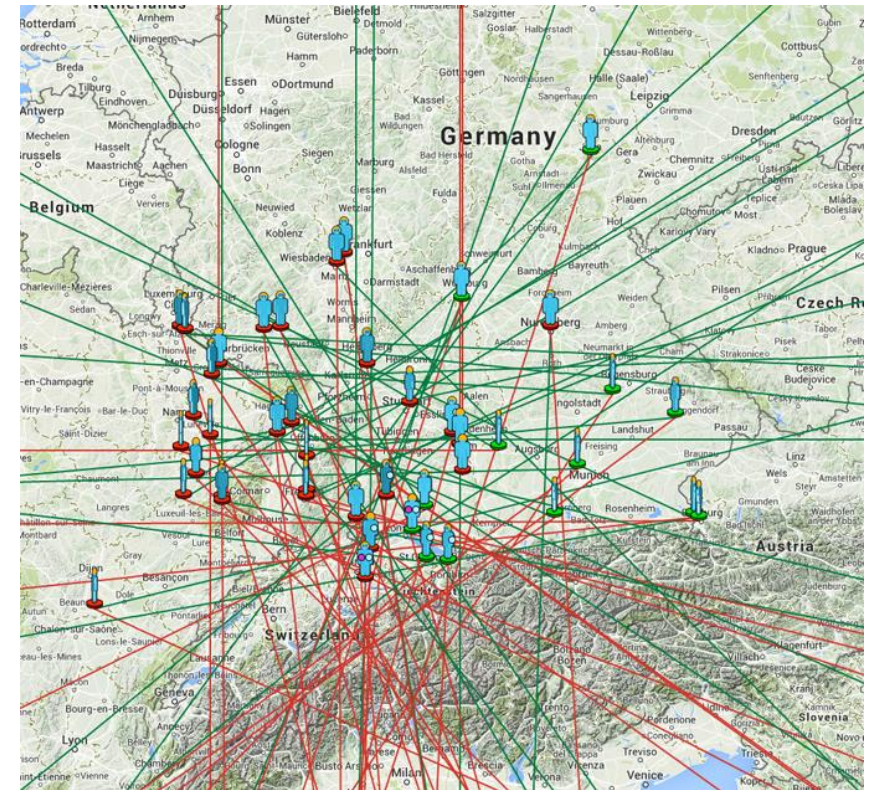
ERGEBNISSE

FEUERKUGEL-AUFZEICHNUNGEN

	Bild	Video	Pfad
13. Juli 2016, 02:21 UT Kurz und bündig			
	VTE SCH	VTE	Map Zoom Daten Animation: xy xz
04. Juli 2016, 23:39 UT Brillante Feuerkugel			
	BOS1, 2 VTE MAU FAL SCH SON BUE1, 2	BOS VTE MAU	Map1, 2, 3, 4 Daten Lichtkurve Animation: xy
18. Mai 2016, 01:06 UT Zum Geniessen			
	VTE BOS SON Radio: VTE BOS SON	VTE BOS Sound: BOS	Map Zoom 1, 2 Daten Lichtkurve Animation: xy xz
20. April 2016, 21:50 UT Klassiker			
	FAL GNO VTE1 VTE2 EGL BAU MAU	FAL VTE1 VTE2	Map Zoom 1, 2 Daten Animation: xy
14. März 2016, 02:27 UT lange Leuchtdauer			
	FAL GNO LOC MAI Radio: ENT	FAL GNO LOC MAI Lichtkurve: MAI	Map Zoom 1, 2 Daten Animation: xy
23. Januar 2016, 21:14 UT über Norditalien			
	LOC VTE SON SCH Radio: BOS VTE MAU ENT SON	VTE Sound: VTE	Map Zoom 1, 2 Daten Animation: xy, xz
	VTE BAU BOS	VTE	Map Zoom Daten

Example: Meteor 15. march 2015

- Observed over Germany and Switzerland
- Video Fredi Bachmann, Olten



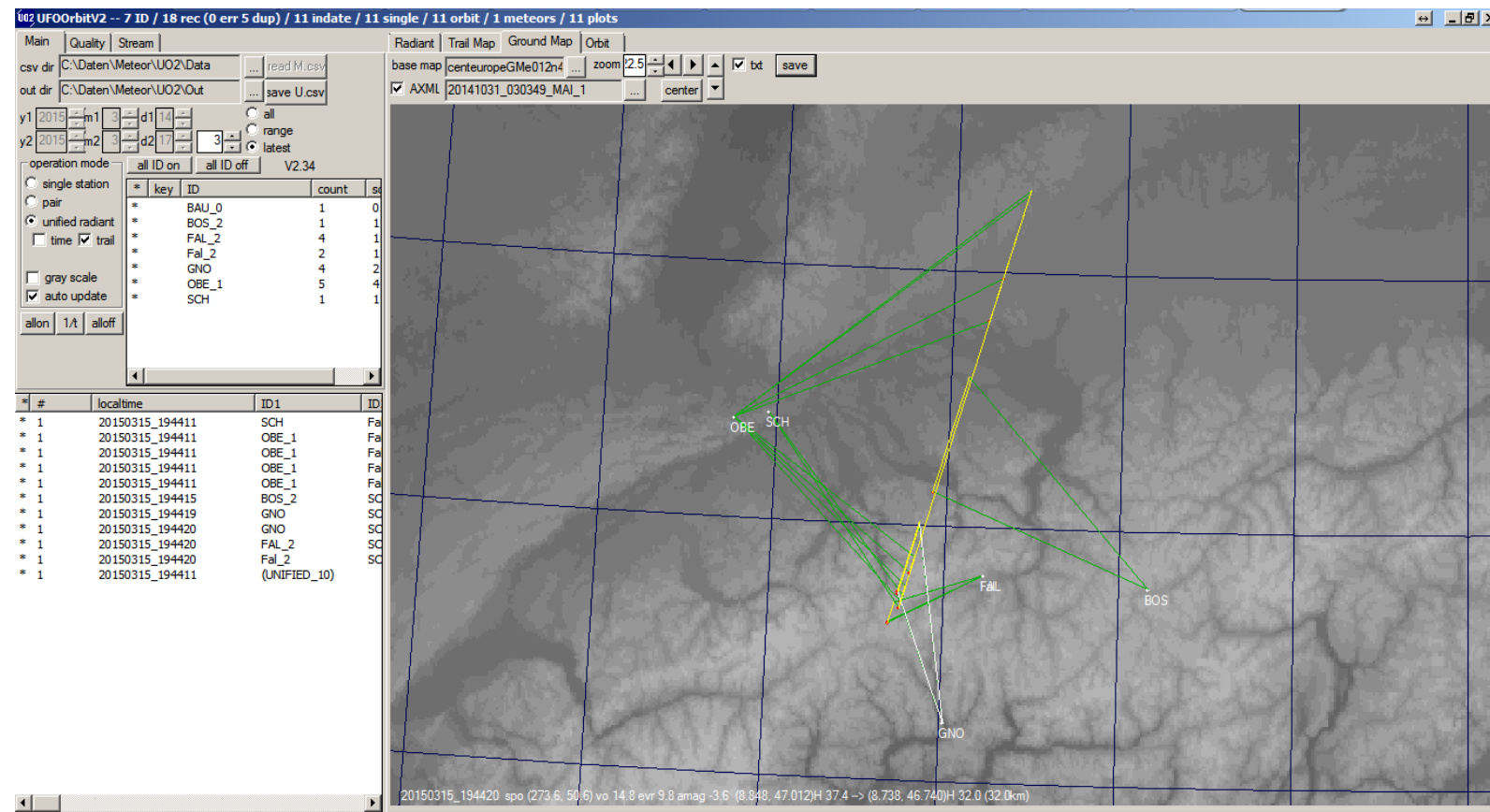
Flight path with triangulation

■ Data from

- Falera
- Oberdorf
- Bos-cha
- Gnosca
- Aarau

■ End point

- $h = 30$ km
- $v = 4$ km/s
- Gotthard

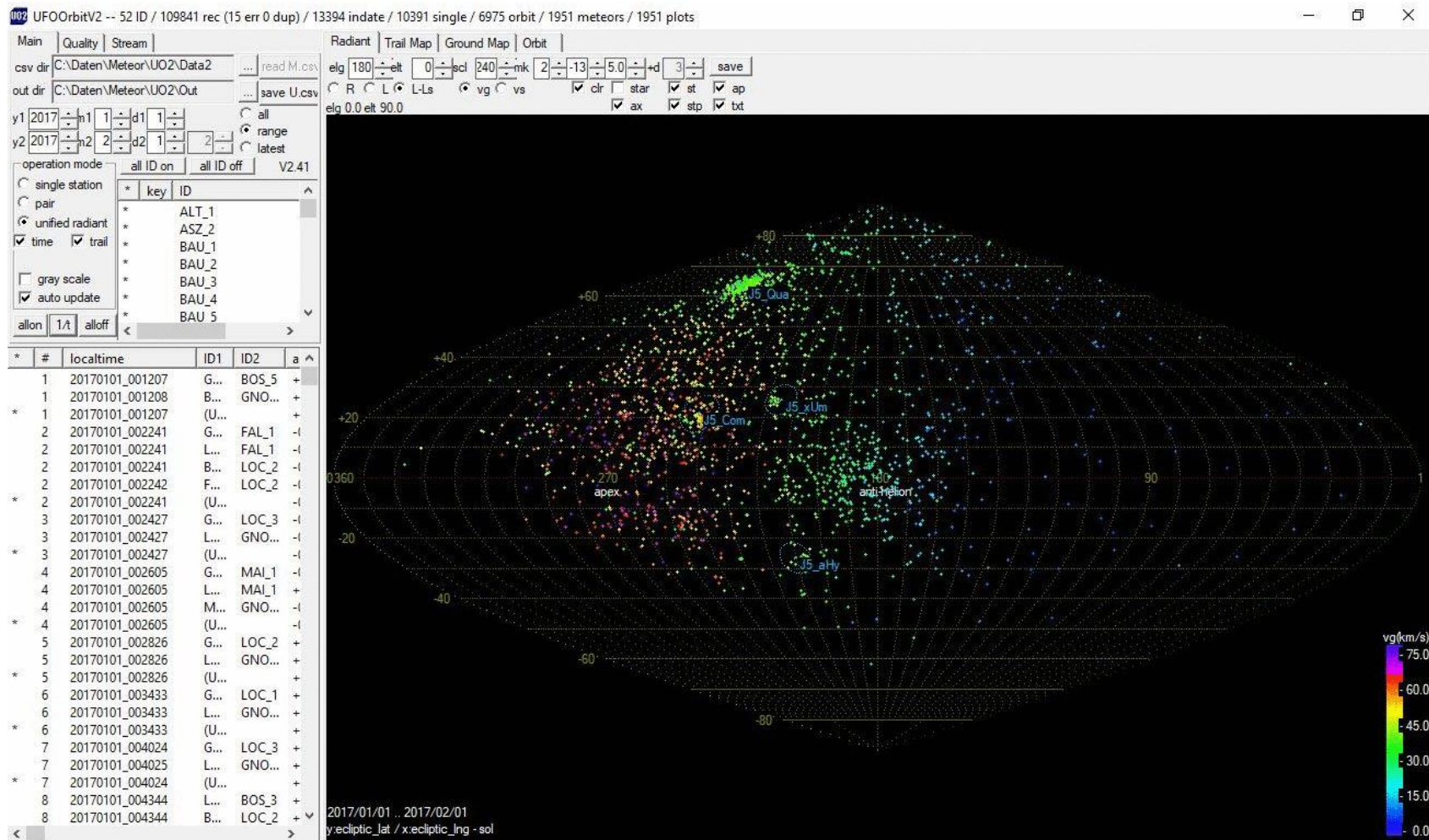


Meteor 15. march 2015

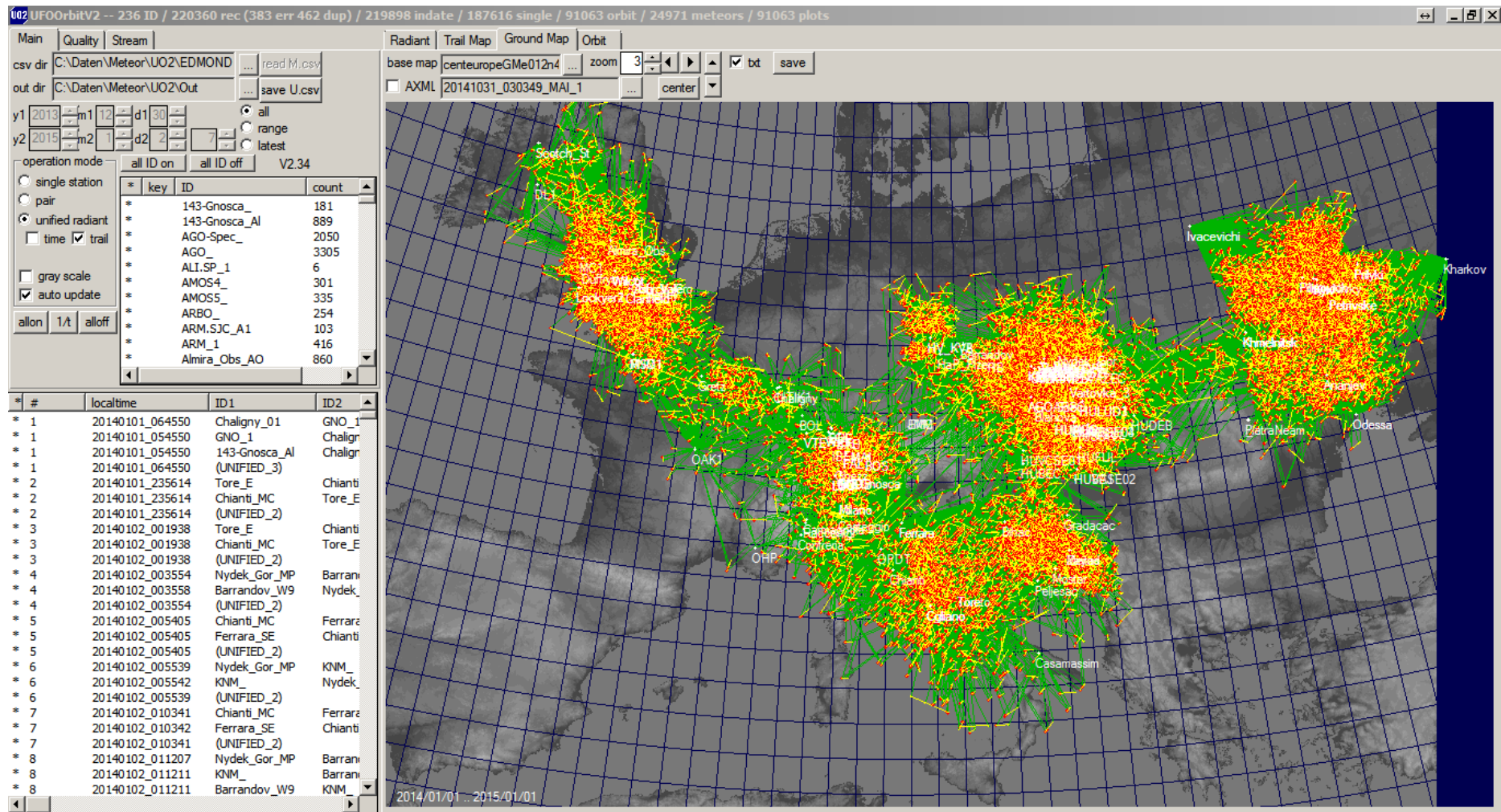
■ search



2017 monthly data



Edmond Database, 2014



Hardware

- Stefano Sposetti
 - Pressure cooker with dome
 - 4x Watec 902H2 ultimate video cameras
 - Wide angle lens f/1.0, 3 – 8 mm
 - Video grabber
 - Intel Core-i7 Computer
- 1 camera to start with

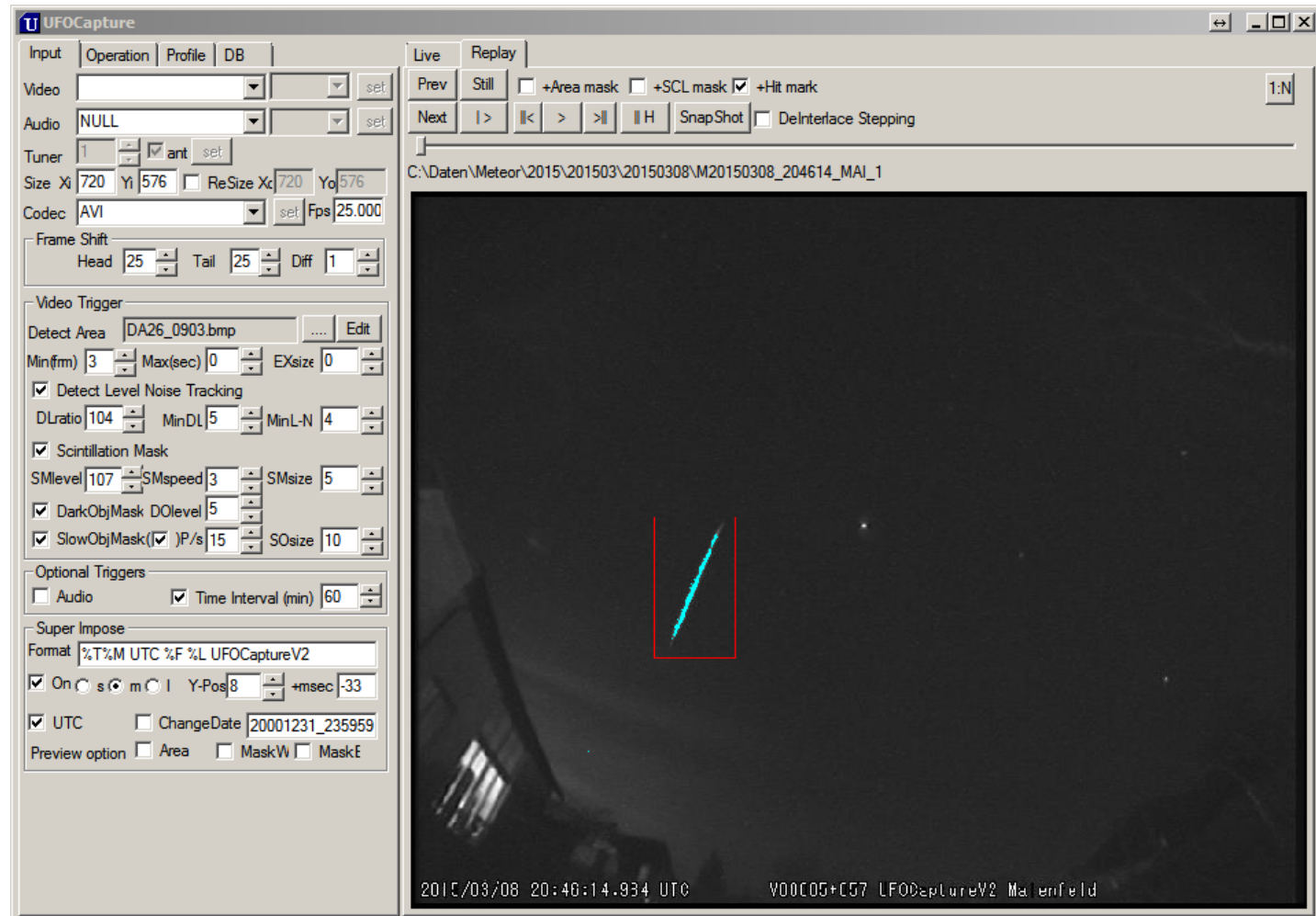


Software

- Metrec (Sirko Molau), special Hardware required
- CAMS (USA)
- FRIPON: French system, especially for bright fireballs
- UFO Tools (Sonotaco, Japan) in Europa widely used, compatible with EDMONDS database
 - UFO Capture: continuous video recording , saving of interesting events with pretrigger (for meteors, lightning, sprites, birds etc.)
 - UFO Analyzer: determination of flight path and velocity, astrometry, star coordinates, magnitude
 - UFO Orbit: determination of trajectory, velocity and orbit from simultaneous observations
 - Data from photographic observations can be used

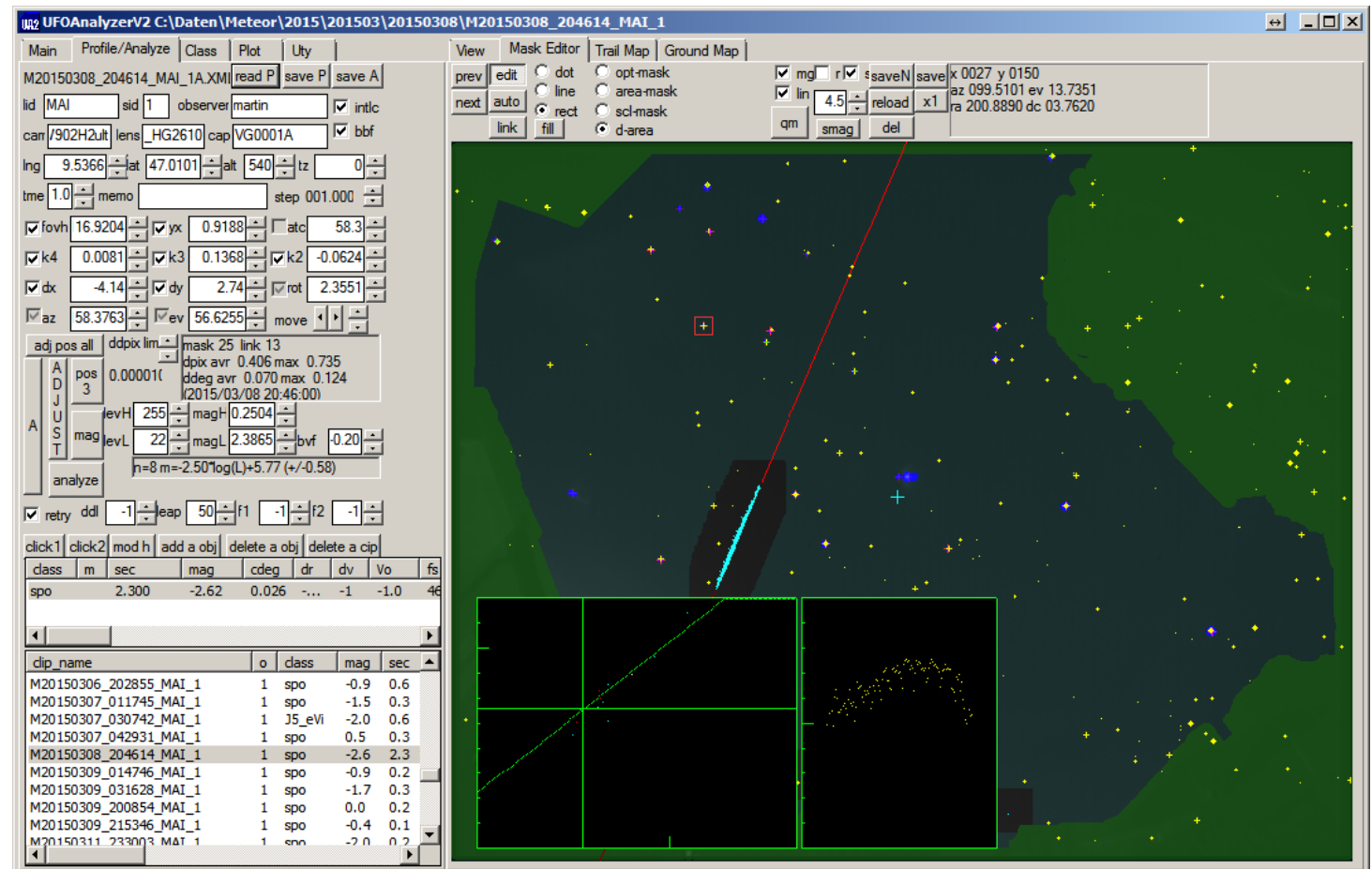
UFO Capture

- Detection
- Capture
- Pretrigger



UFO Analyze

- Astrometry
- Flight path
- Magnitude
- Velocity

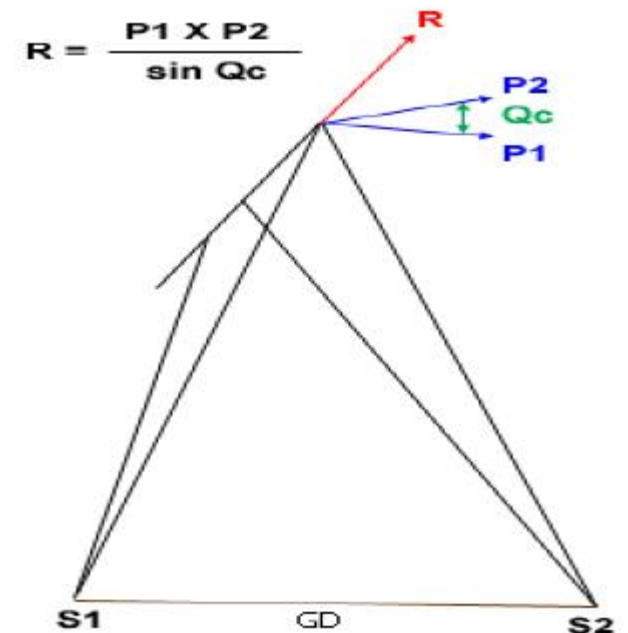


Orbit calculation with UFO Orbit

■ Terrestrial observer

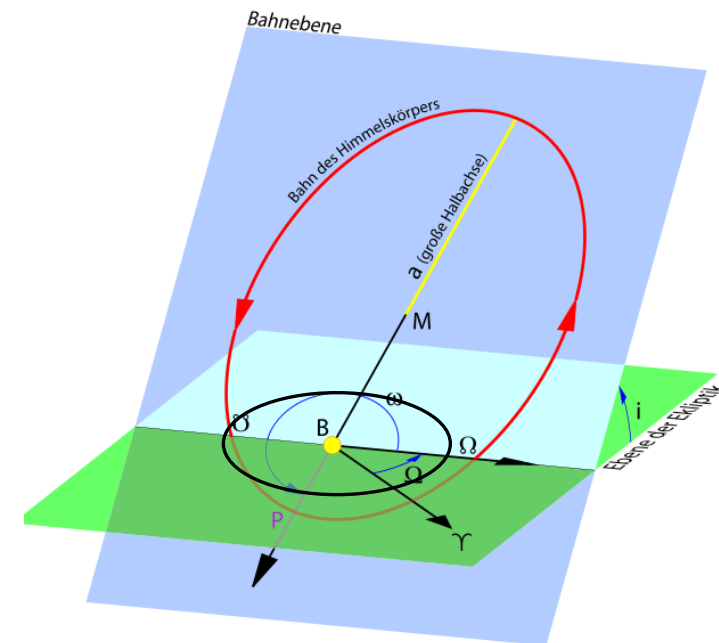
- Radiant and velocity observed
- Correction for deceleration in atmosphere
- Correction for earth rotation:
 $0.456 \cdot \cos(\varphi)$ km/sec, rund 300 m/sec $\rightarrow v_g$
- Velocity correction for gravitational potential:
 $v_0^2 = v_g^2 - 125.4$
(v in km/sec)
- Correction zenital attraction
 $\Delta z = 2 \cdot \text{ARCTAN}((v_g - v_0)/(v_g + v_0) \cdot \text{TAN}(z/2))$

■ \rightarrow velocity relative to earth outside gravitational potential



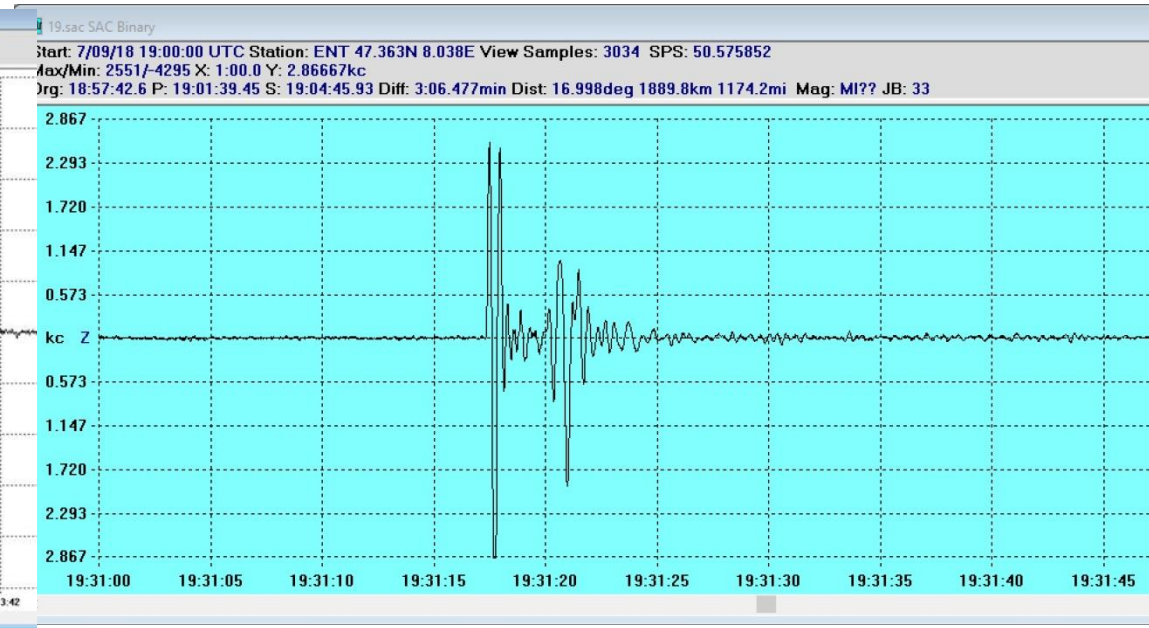
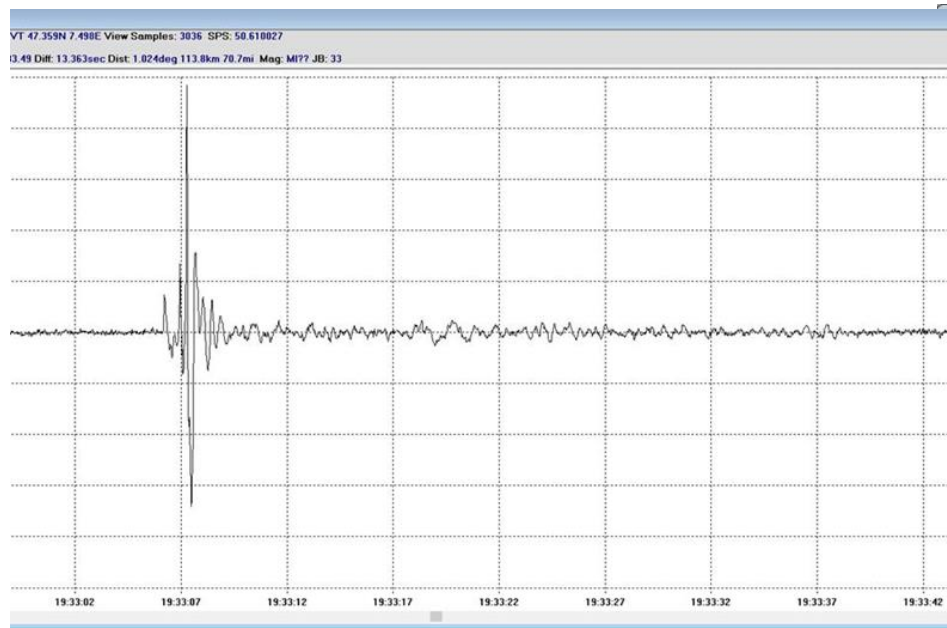
Orbit calculation with UFO Orbit 2

- Velocity relative to earth
 - (Vector-) addition earth orbit velocity (approx. 30 km/sec) →
 - Heliocentric meteor velocity (x, y, z, v_x, v_y, v_z)
 - From distance (1AE) and velocity → semi major axis a
 - orbit time of revolution (3rd law of Kepler)
 - Vector v and r (Vector sun – meteor) → angular momentum (→ eccentricity e),
 - Orbit plane (inclination i , node Ω)
 - → position of perihel ω , T perihel



Other activities in meteor observations

- Radio detection
- Infrasound measurements
- Detection of transient luminous events (TLE) in the upper atmosphere



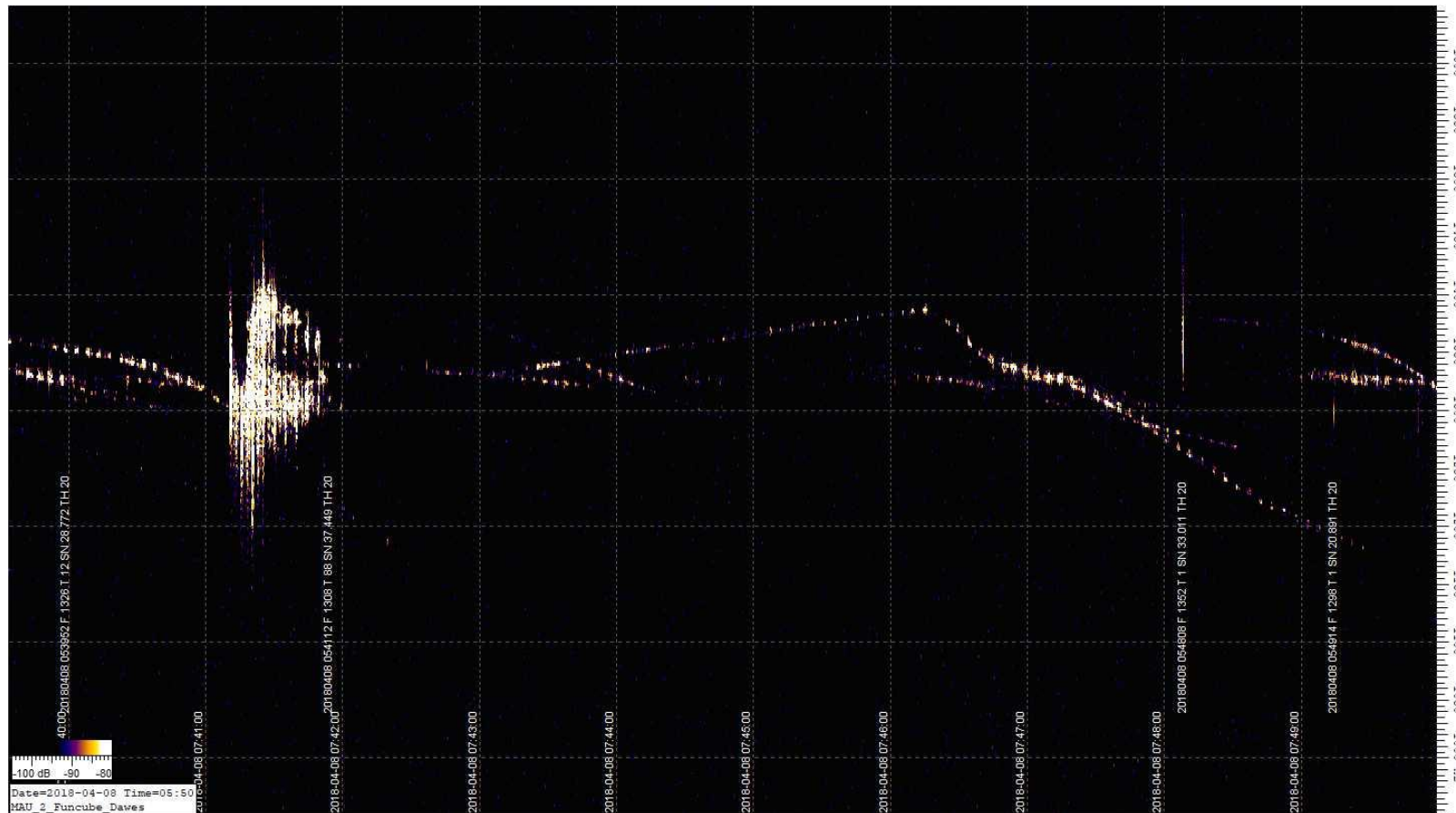
TLE, sprite

- First detection at Maienfeld



Radio detection, echo from Graves 143.05 MHz

- Daylight event R20180408_054130 UT



Fachgruppe Meteorastronomie



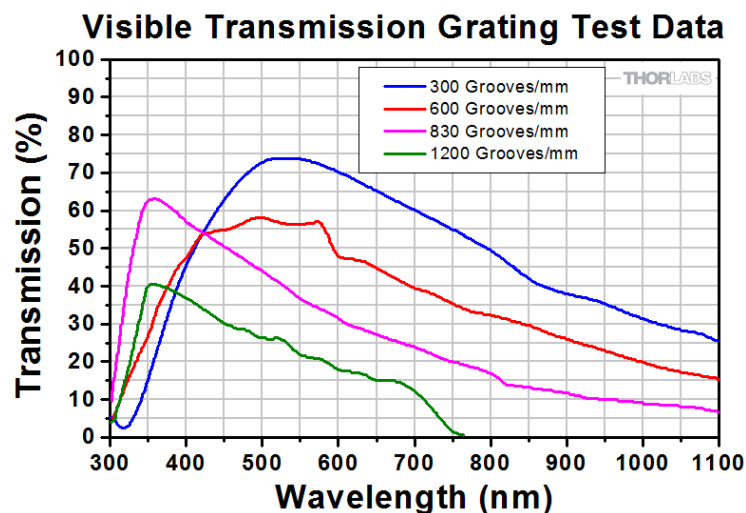
<http://www.eso.org/public/images/potw1414a/>

Meteor spectroscopy

- Hard- Software
- Wavelength calibration
- Spectrum extraction
- Instrument response
- Summary

Hardware

- Watec 902H2 ult. Computar HG2610AFCS-HSP F/1 2.6mm fl
- 902H2 ultimate (spectroscopy) Tamron 12VG412ASIR F/1.2, $\approx 7\text{mm}$ fl
- 2nd camera with transmission grating for spectroscopy
Thorlabs
300 L/mm \rightarrow 600 L/mm



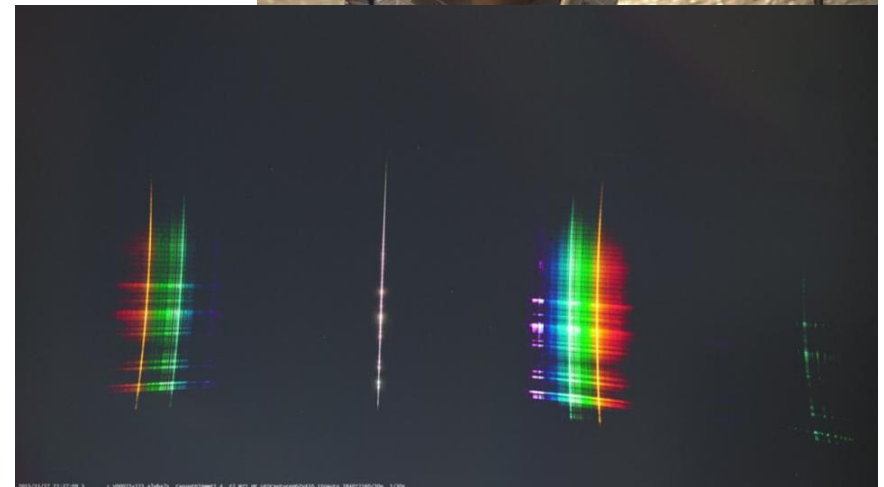
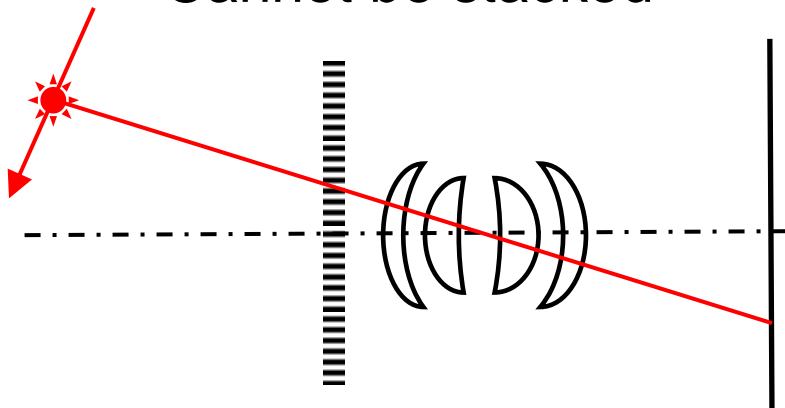
OHP 2017

- Mobile setup



Starting point

- Camera with wide angle lens
- Transmission grating
 - mounted **perpendicular** to optical axis!
- Problem:
 - Moving meteor
 - Curved spectra with nonlinear dispersion
 - Cannot be stacked



Spectrograph, theory

- Video camera with transmission grating in front of lens

- Grating equation:

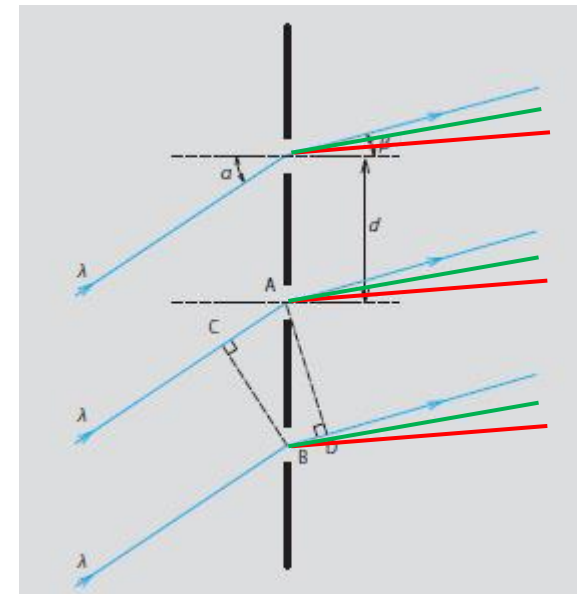
$$- m * \lambda * G = (\sin \alpha - \sin \beta) * \cos \gamma$$

- m: grating order, G: grating lines / mm
- λ : wavelength
- α , β : angle of incidence, transmitted beam
- γ : cross, out of plane angle

- Inverse dispersion per pixel:

$$d\lambda/dx = (\cos \beta \cos \gamma) / (m * G * f) * p \quad (p: \text{pixel size})$$

- Example: $f = 7 \text{ mm}$, $p = 8.6 \mu\text{m}$, $G: 300\text{L/mm}$ $\beta = 0 \rightarrow d\lambda/dx = 41\text{\AA/pixel}$



Vector theory wavelength calibration

- Grating in front of lens perpendicular to optical (z-)axis

- Unit vector (A B C) for incident direction

- Diffracted beam

$$A' = A + m\lambda G \quad (x\text{-axis})$$

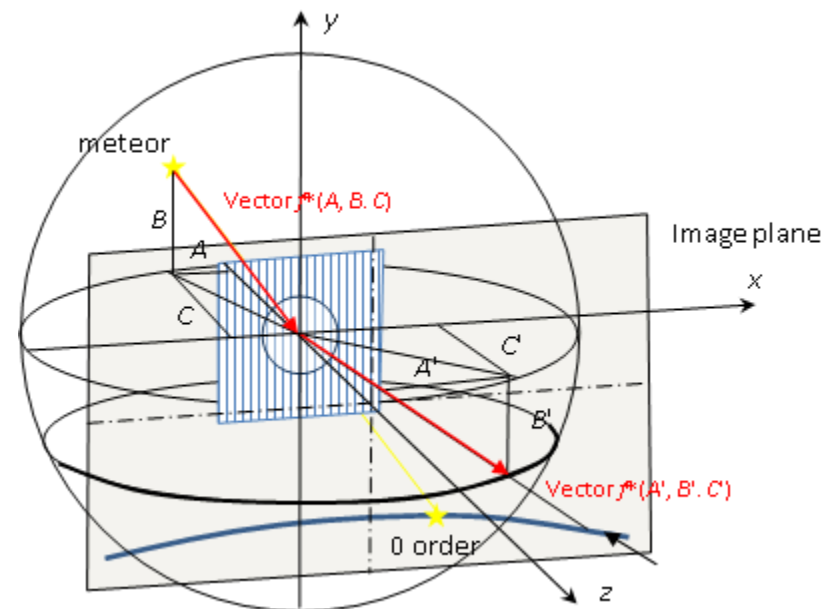
$$B' = B \quad (y\text{-axis})$$

$$C' = \text{sqrt}(1 - A'^2 - B'^2)$$

- Spectrum on CCD plane

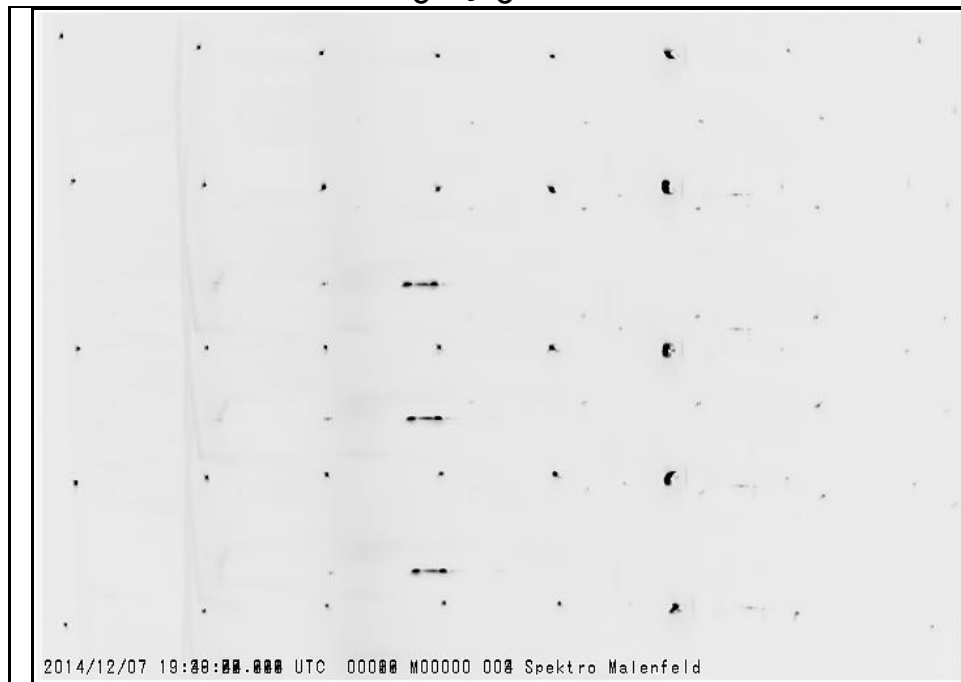
- Nonlinear dispersion
- Hyperbolic curvature

- Spectrum straight linear in A',B'

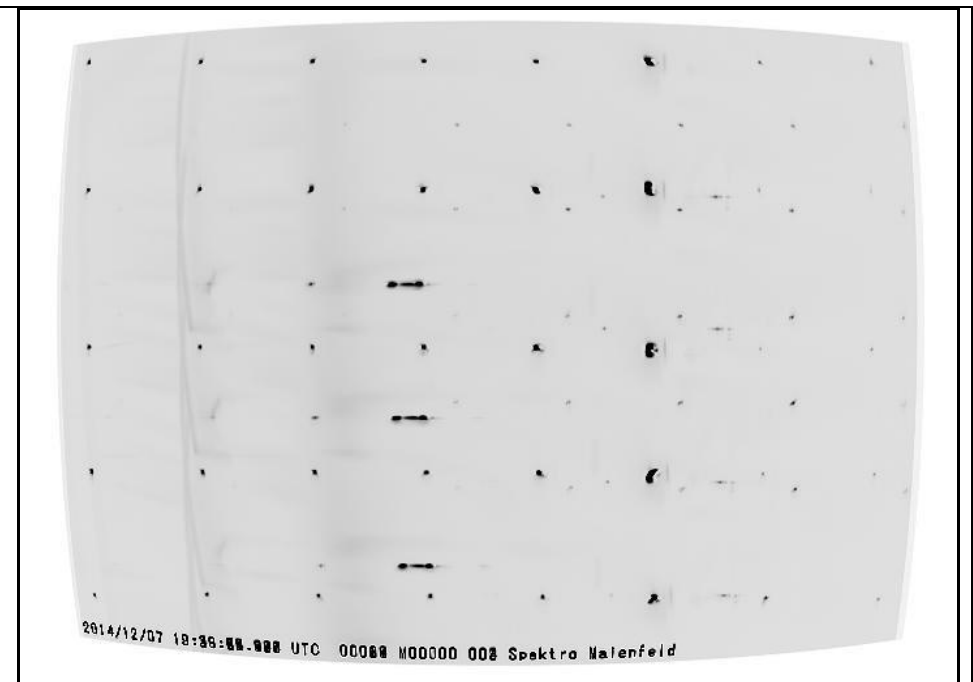


Calibration spectrum HeNe laser

- HeNe laser $\lambda = 633 \text{ nm}$, $f = 4 \text{ mm}$
- Fit with polynom $r = r' * [1 + 3.94E-07*r'^2 + 2.01E-12*r'^4]$
- Fit center x_0, y_0

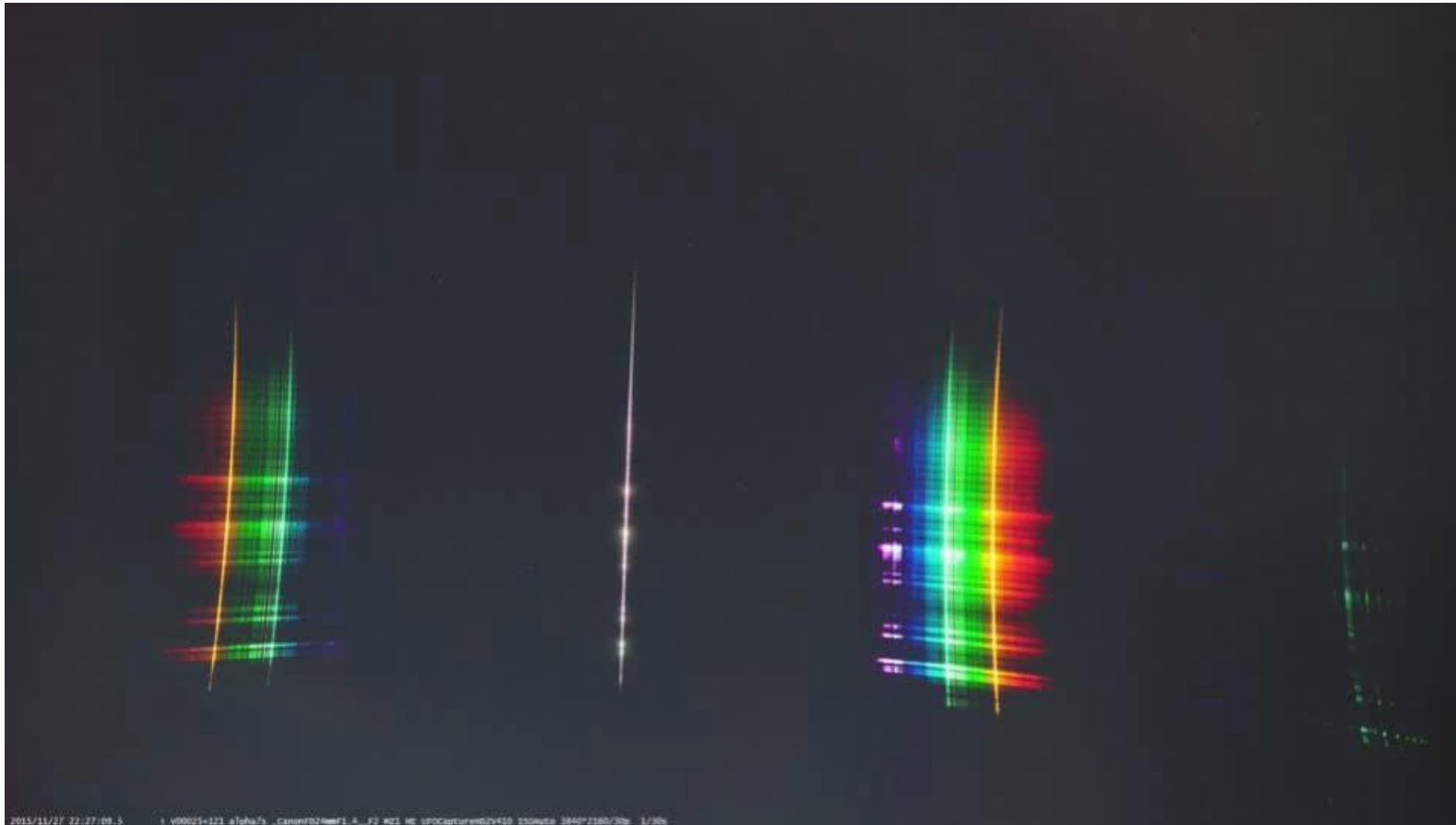


Composite spectra original

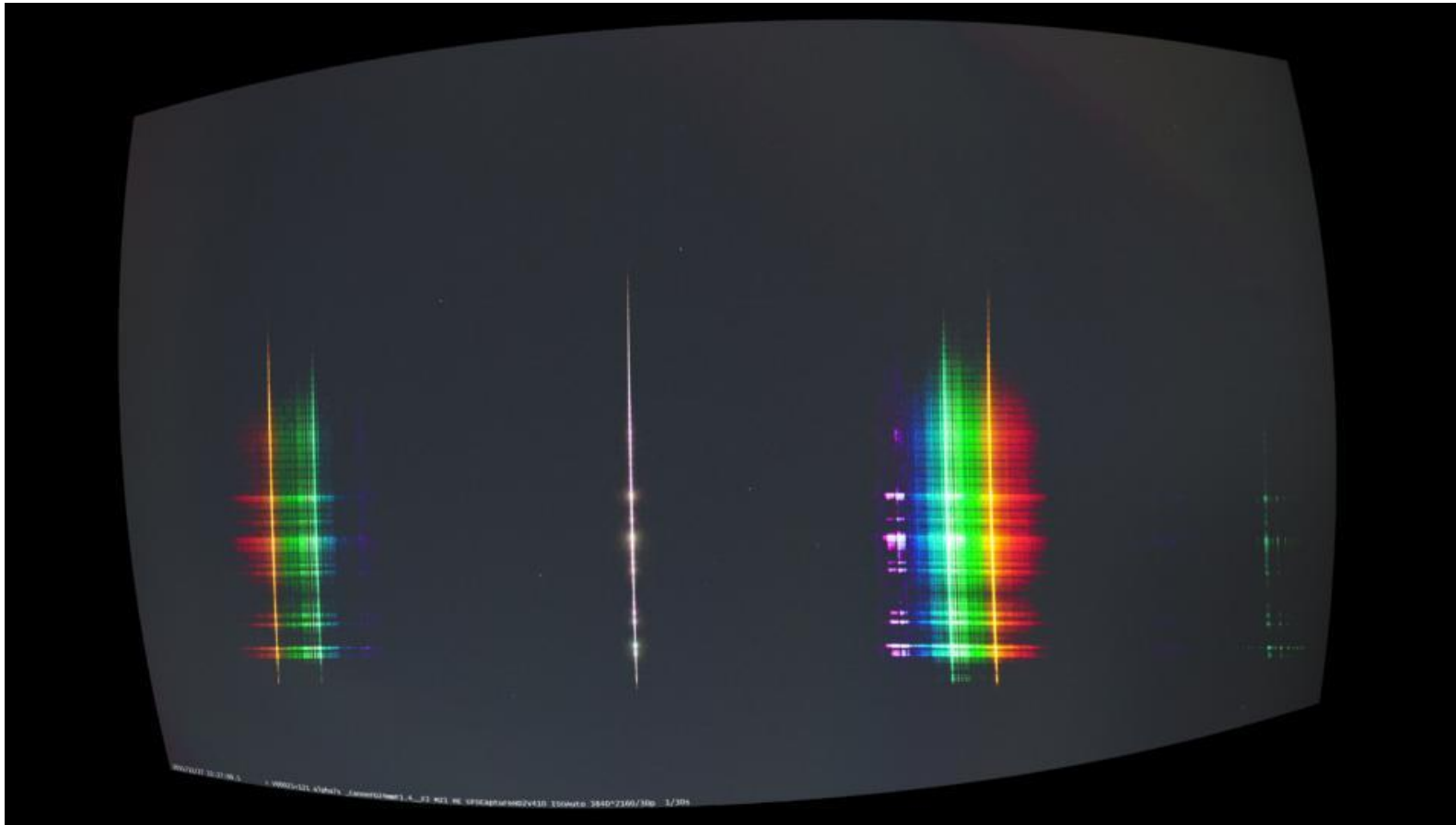


After applying transformation

Orthographic transformation, original



Orthographic transformation, result



Orthographic transformation, result

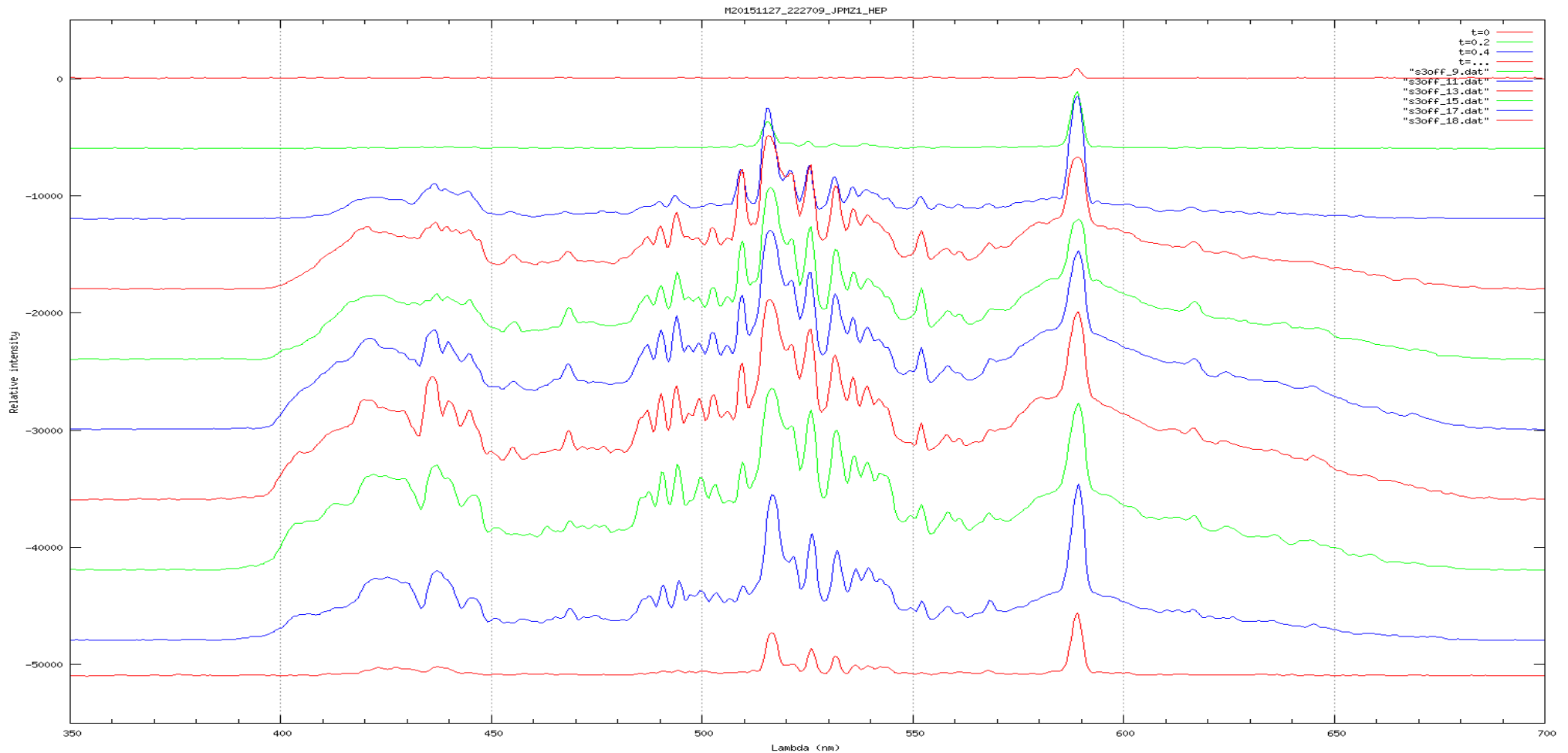
- Frames converted to b/w, linearized, registered, M20151127_222709



- color



Extraction of spectra



Full processing

- Wavelength calibration ✓

- Flux calibration

Correct for:

- Background subtraction!
- Vignetting, field of view
- Correction for image transformation
 - Apply image transformation
 - Extract spectrum, calibrate wavelength

} flat field correction
in pre-processing

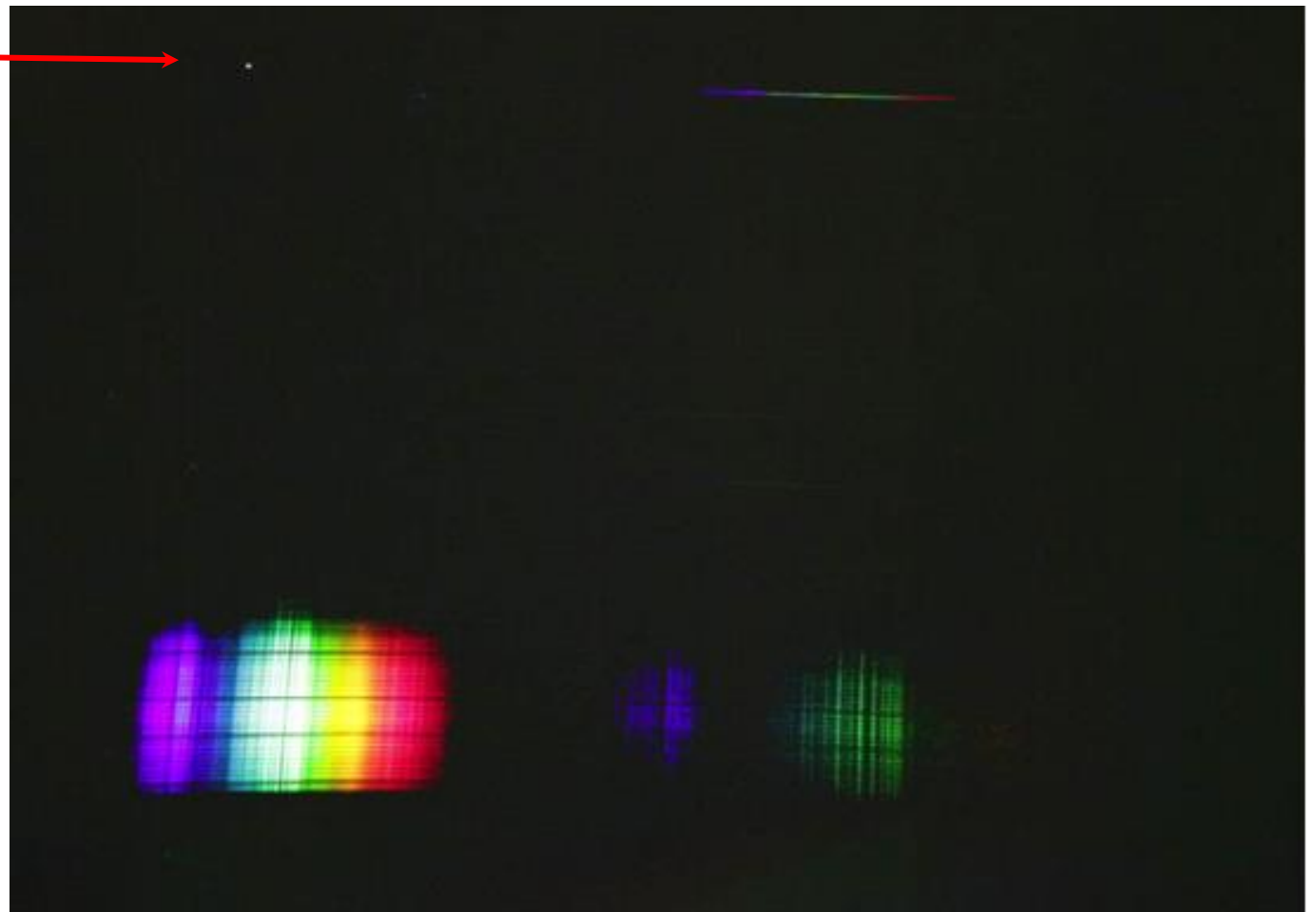
- Instrument response

- Grating efficiency
- Camera spectral sensitivity (lens, CCD)
- Atmospheric transmittance

} instrument response

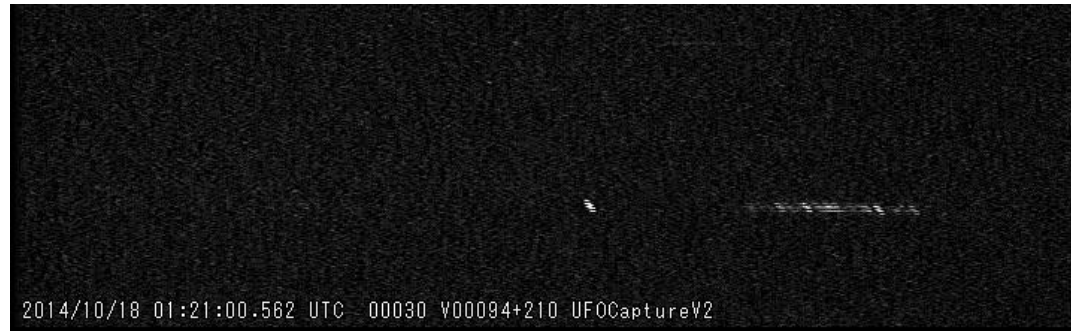
M20160107_225526_JPMZ1_HDP, iron meteor

- Sirius →
- Koji Maeda
- Canon EOS6D
Sigma 35mm/f1.4
- zero order missing



Preprocessing

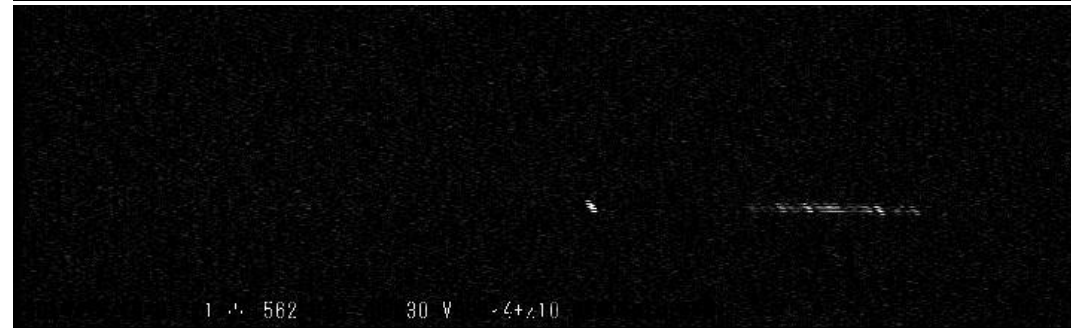
- Extract image (i30)



- Background ADD_MEAN < I1 ... I20 >



- Subtraction SUB2



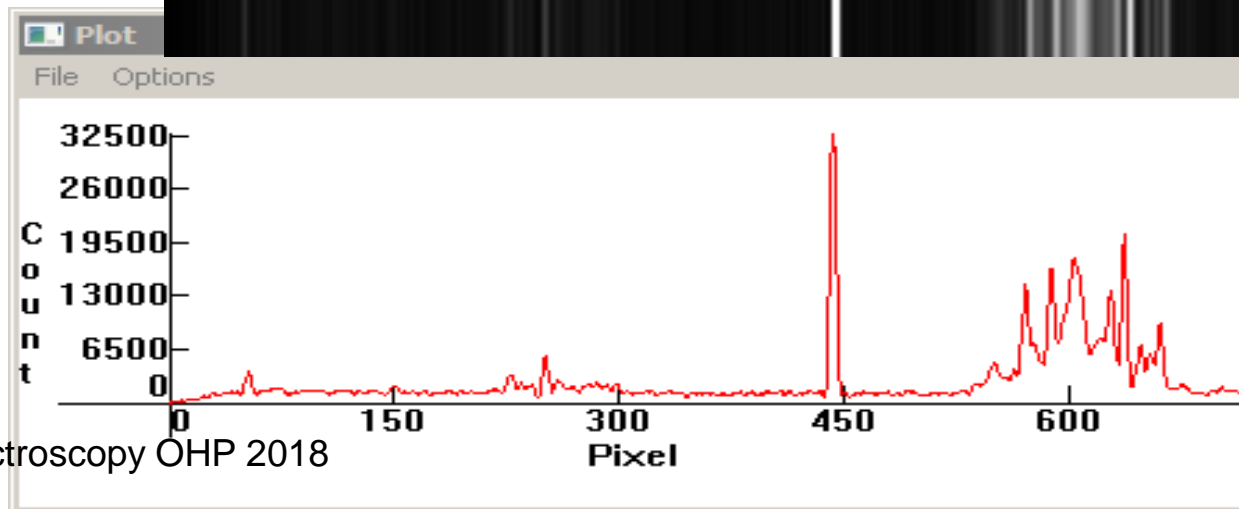
Processing 2

- Add
TRANS, ADD
ib34-ib40

- Slant 28 30

- L_ADD

- L_PLOT
save file.dat



Background subtraction with IRIS

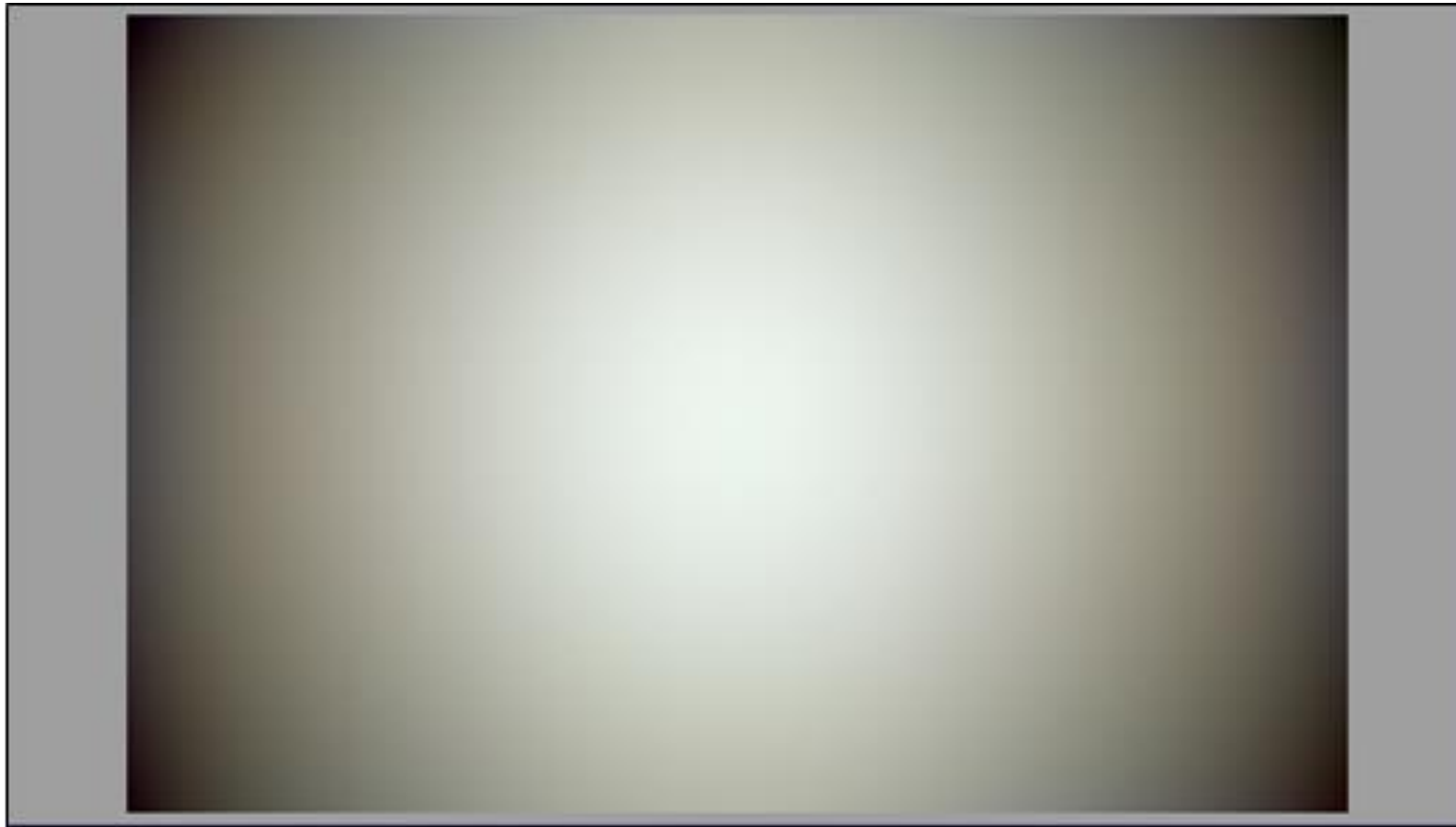
- Pretrigger → no meteor visible, used for background
 - M_BACK.PGM

```
CONVERTBMP24bw $1 @b $3
*CONVERTBMP24 $1 @b $3
ADD2 @b $3
MULT $4
SAVE $2
REMOVE@
```
- Background subtraction: M_DARK.PGM
 - SETBASE \$4

```
CONVERTBMP24BW $1 @m $5
*CONVERTBMP24 $1 @m $5
SUB2 @m $2 $3 0 $5
REMOVE@
SETBASE 1
```

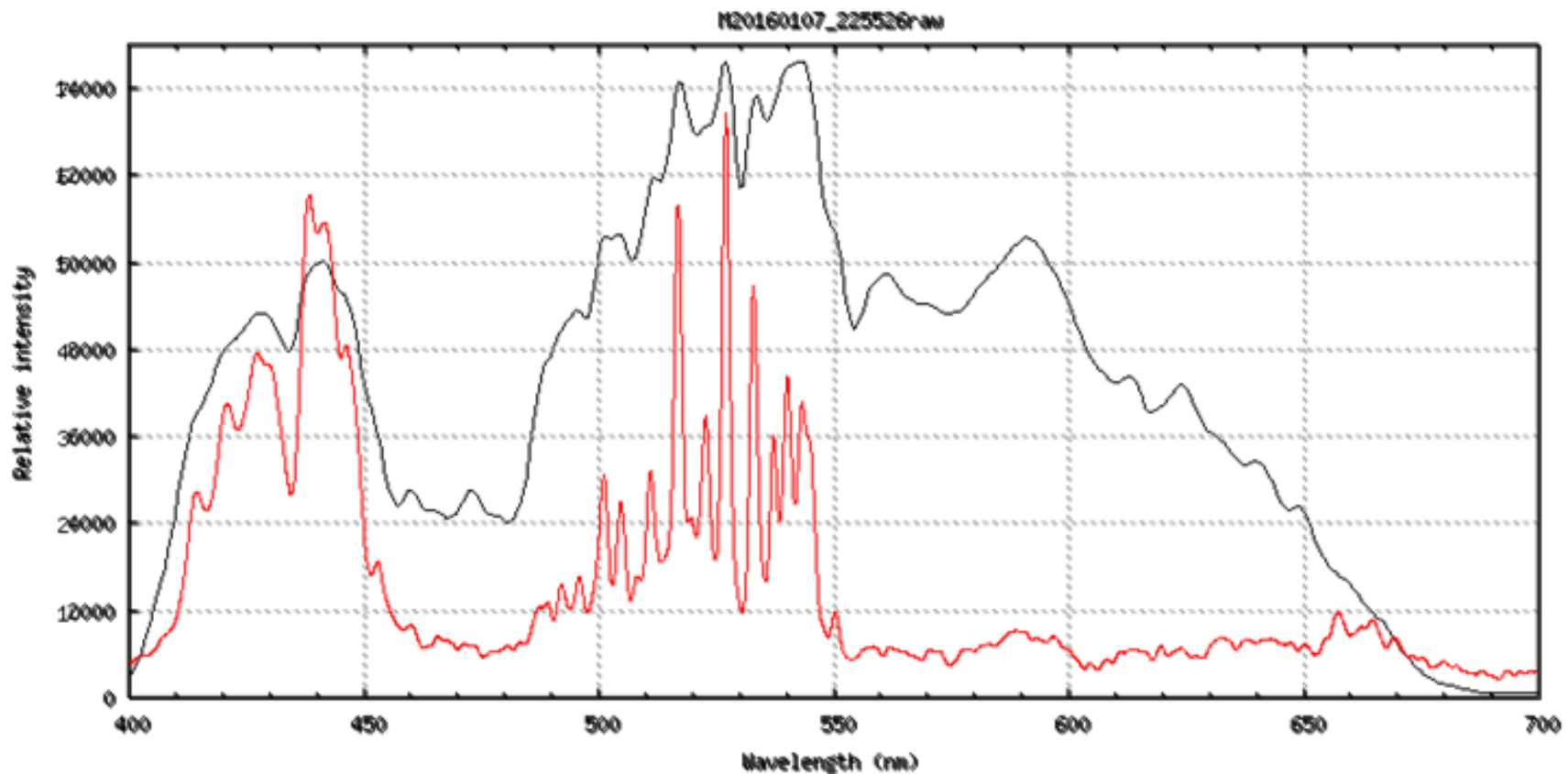
Flat field

- Border not illuminated



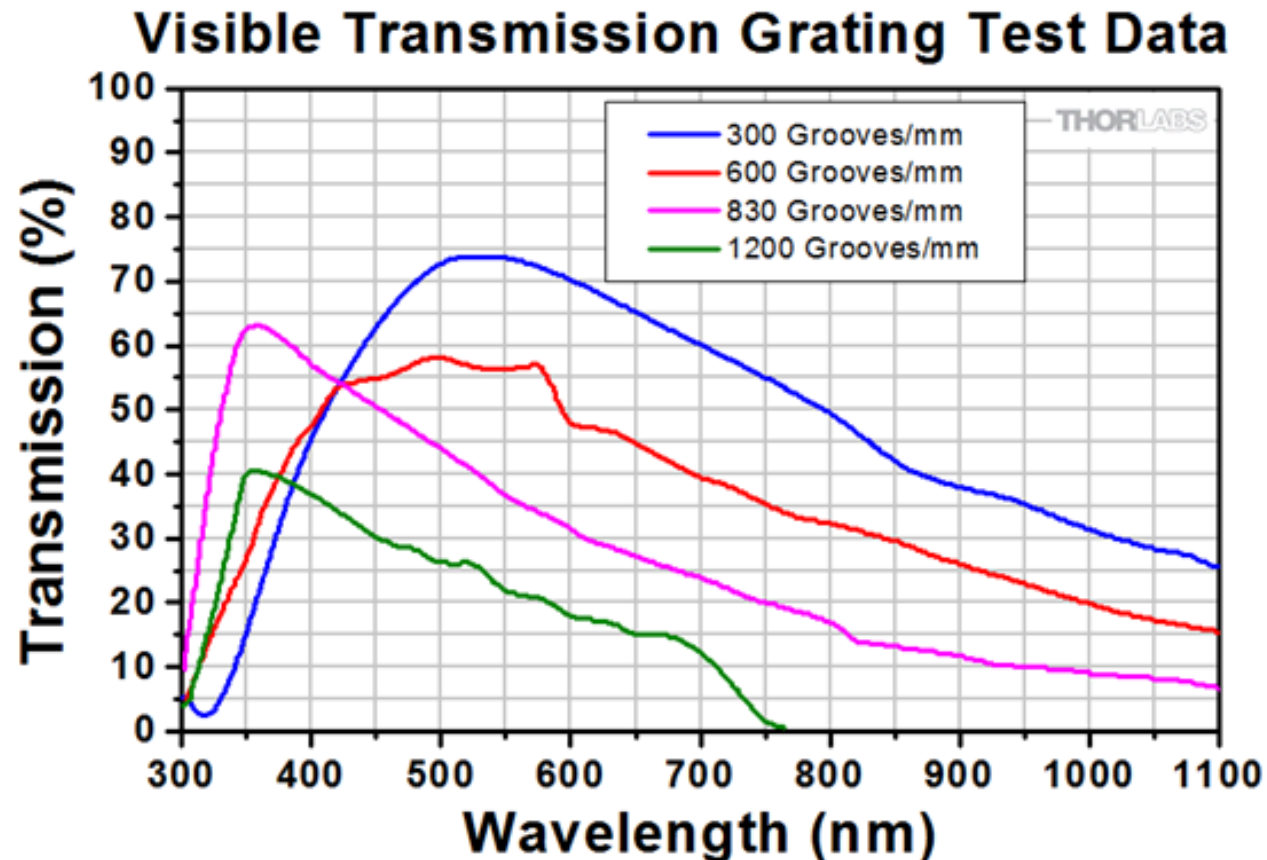
Saturation

- 1st and 2nd order on same wavelength scale



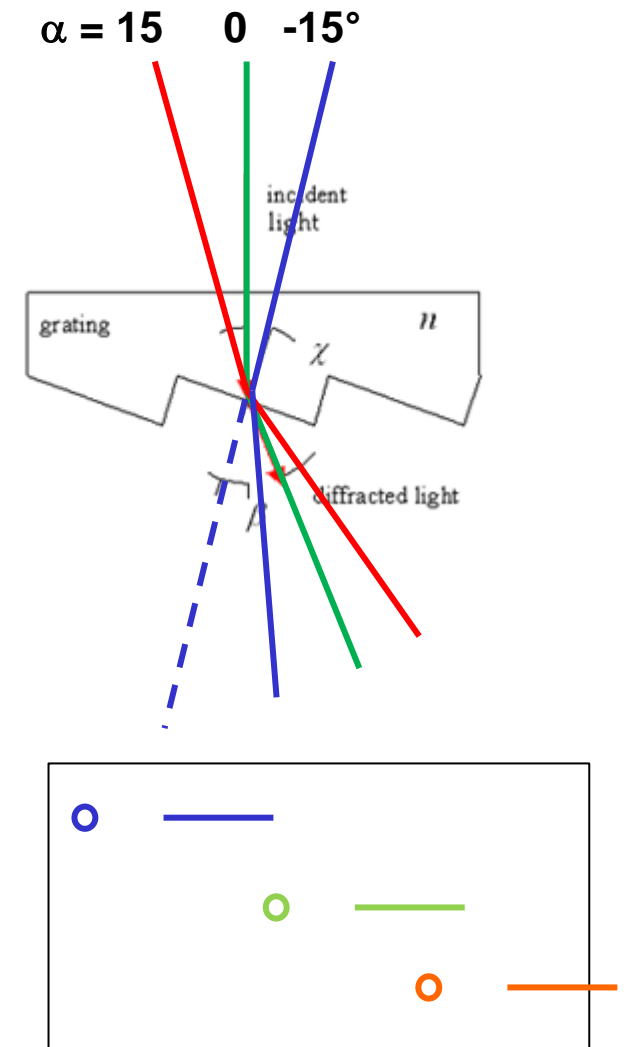
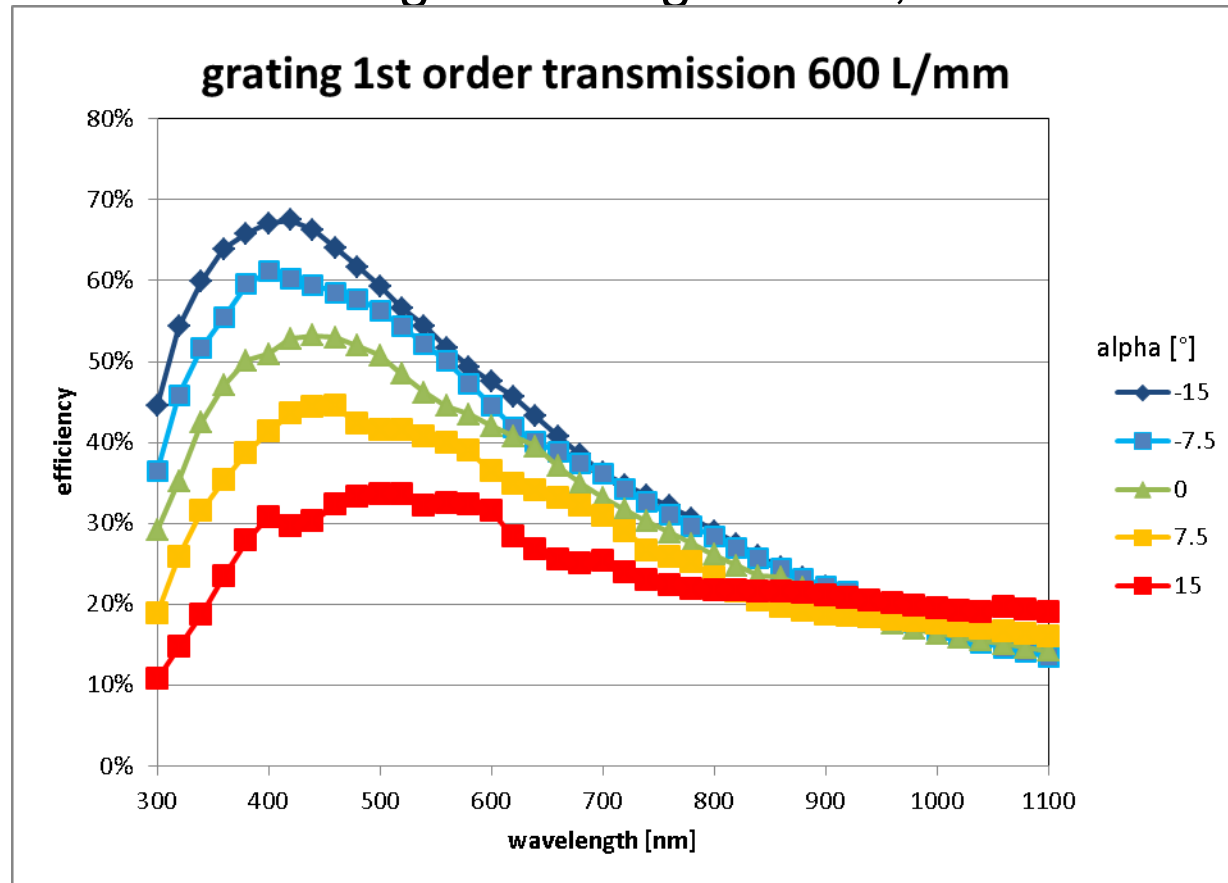
Grating efficiency

- Depends on grooves/mm and groove angle
- Thorlabs:



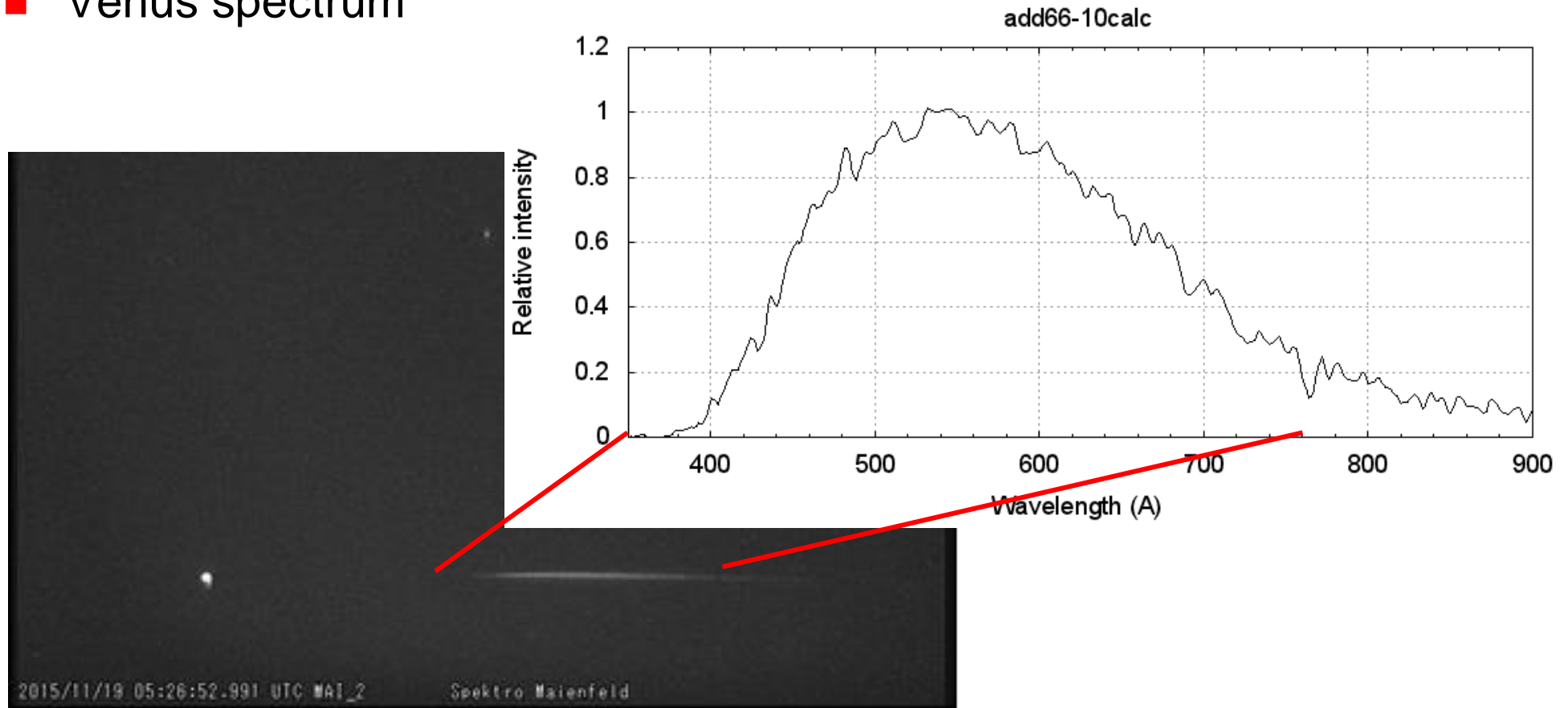
Grating efficiency 2

- Grating efficiency dependent on incident angle!
Calculated for groove angle 28.7° , $n = 1.52$



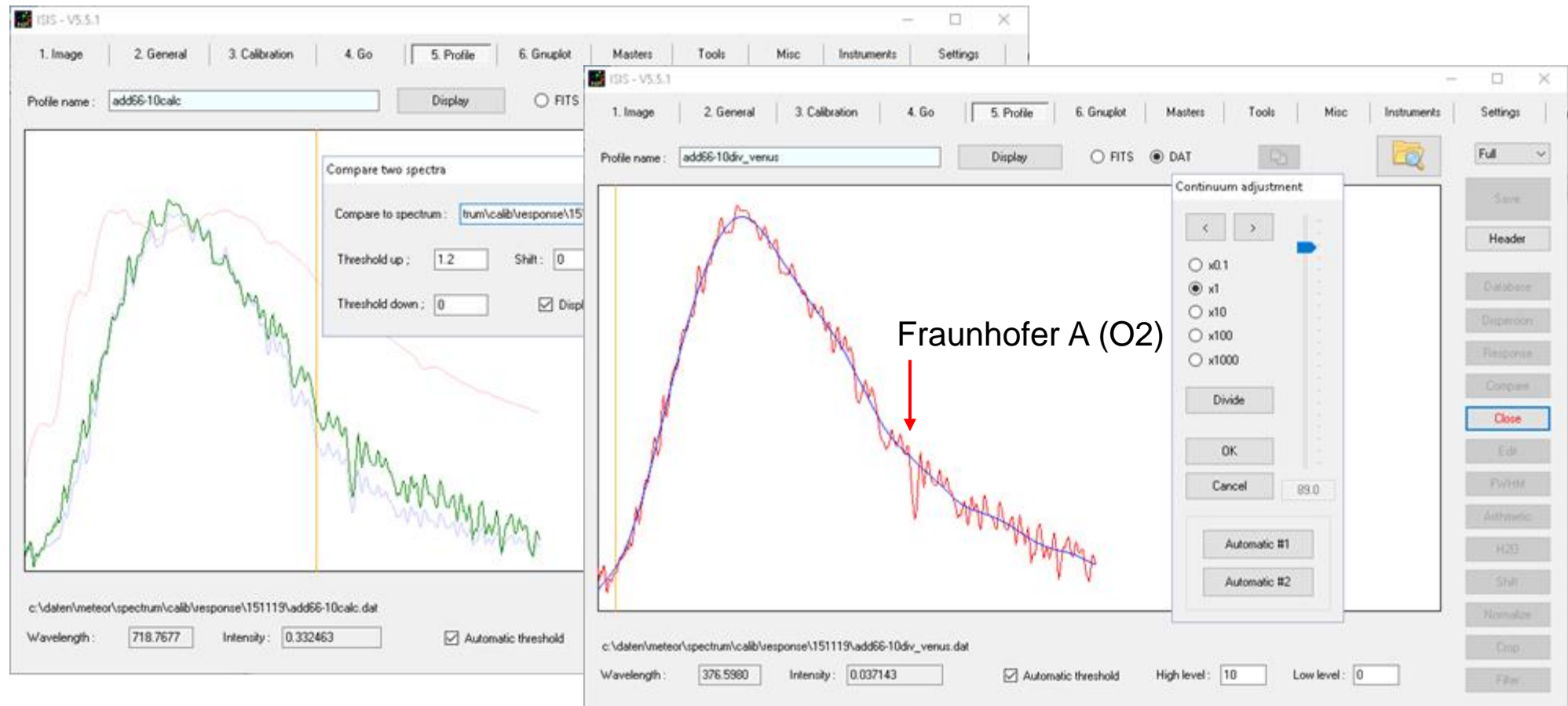
Measured spectrum, Watec 902H2ultimate

- Venus spectrum



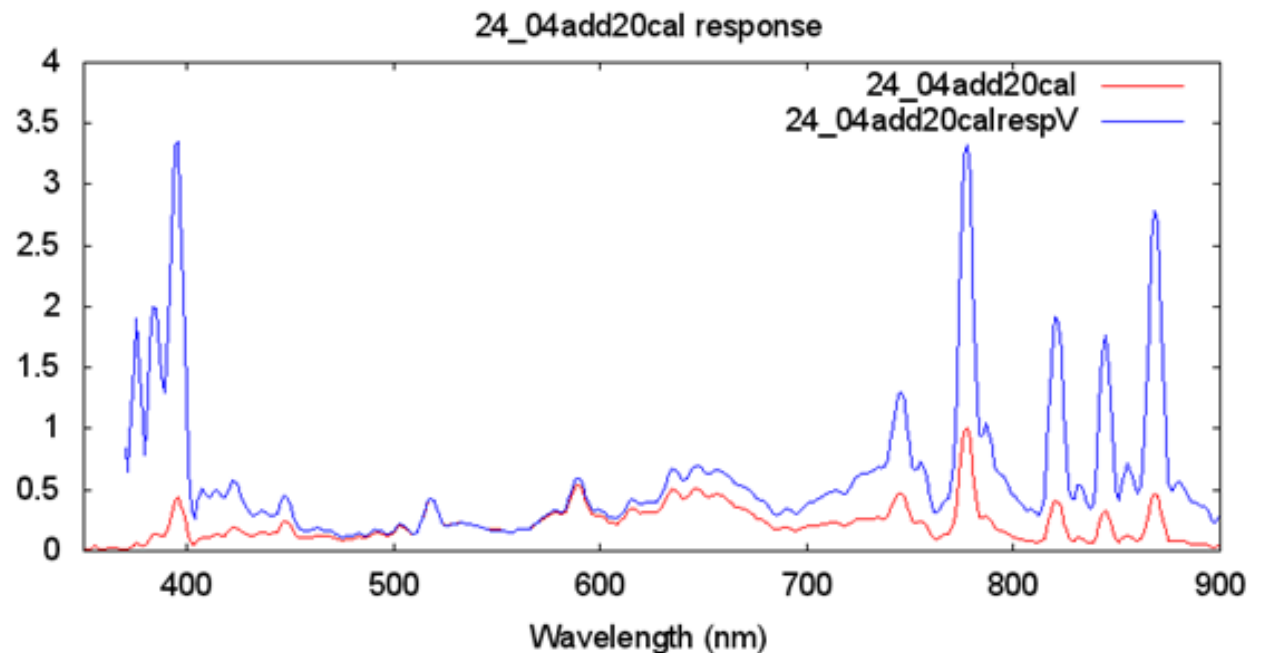
Instrument response

- Measured spectrum / reference spectrum



Instrument response

- Spectrum of known object (Venus)
- Instrument response = measured spectrum / reference spectrum
- Meteor spectrum, wavelength calibrated
- Corrected spectrum = meteor spectrum / instrument response



Characteristic meteor spectra

- Catalogue of meteor spectra: V. Vojacek et. Al.
<http://adsabs.harvard.edu/abs/2015A%26A...580A..67V>

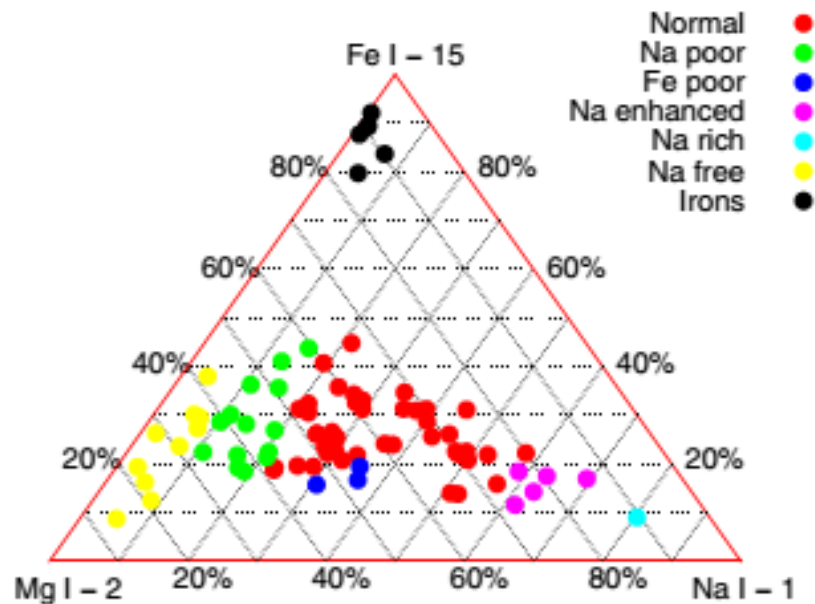


Fig. 8. Classification of meteor spectra. The ternary graph of the Mg I (2), Na I (1), and Fe I (15) multiplet relative intensities. Every group of meteoroids is represented with a different symbol.

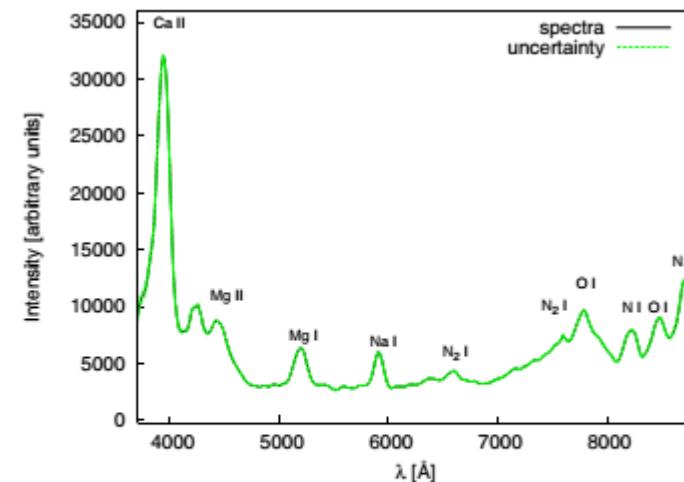
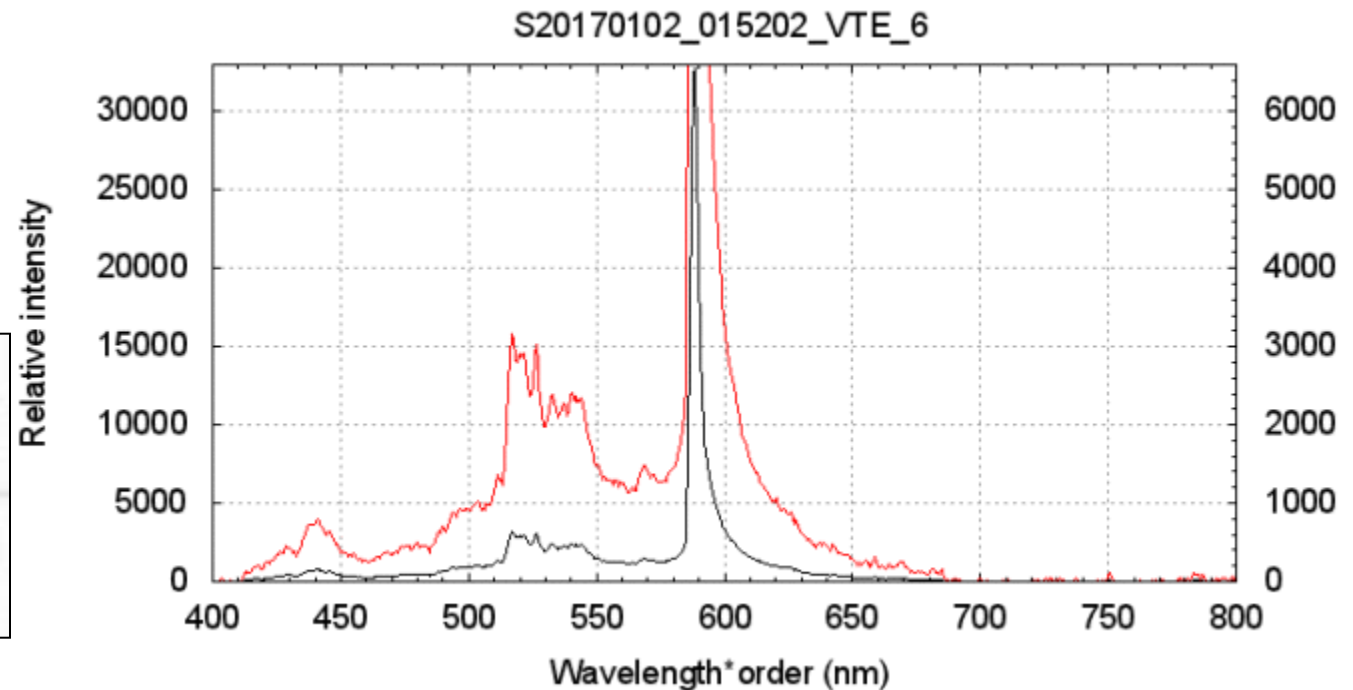
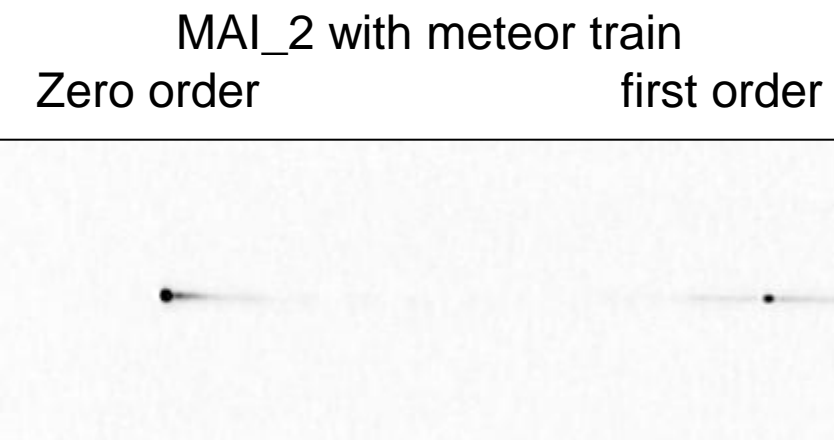


Fig. 5. Spectrum SX1837 of a bright Perseid. The meteor had a maximum brightness of -9.2 mag. Because the spectra were oversaturated on the video frames around the brightness maximum, one frame of the sequence was chosen. The brightness of the meteor in this frame was -7.5 mag.

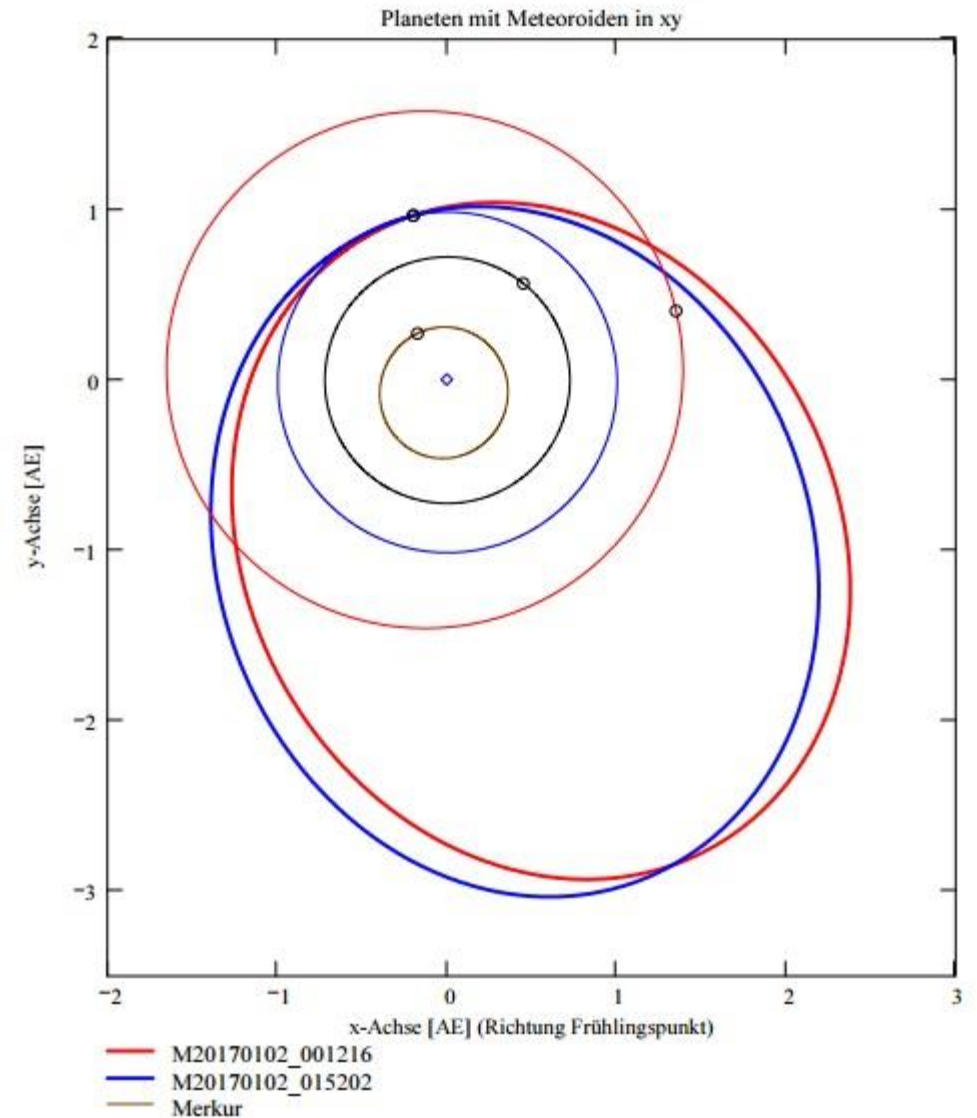
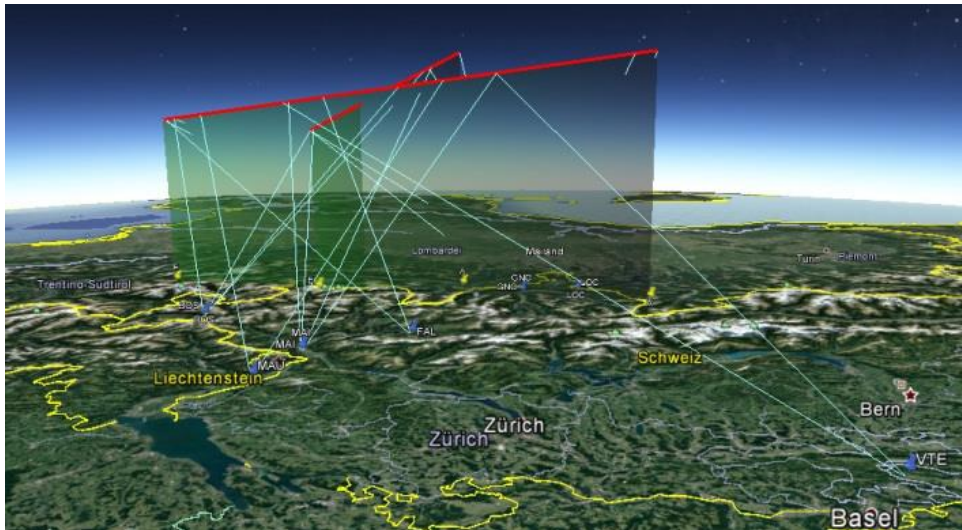
Slow meteor

- <https://meteorspectroscopy.org/blog/two-slow-meteors-with-spectra/>
- Strong Na-lines



2 slow meteors

- $v_0 = 13 - 14$ km/sec
- 20170102_001216
20170102_015158
- Similar orbits



Conclusion

- Grating mounted perpendicular to camera axis
- Orthographic image transformation gives linear spectra!
- Precise flux calibration depends on many factors, approximations used
- Looking for low cost, sensitive, high resolution, high dynamic range video camera
- Full format colour camera (e.g. Sony)
 - + Color → easy interpretation
 - + Orders can be separated
 - + High resolution
 - Bayer matrix lower sensitivity
 - Difficult to analyse (Instr. Resp.)
 - cost
- Video camera (e.g. Watec)
 - + High sensitivity
 - + Spectral range
 - + Low cost
 - Small field of view or
 - Low spectral resolution
 - Overlapping orders
 - Low dynamic range

Spectrum recording and processing software

- UFO Capture for trigger and record video
(http://sonotaco.com/e_index.html)
- IRIS (<http://www.astrosurf.com/buil/us/iris/iris.htm>)
astronomical image processing and spectroscopy software
http://www.astrosurf.com/buil/iris/nav_pane/CommandsFrame.html
- ISIS (http://www.astrosurf.com/buil/isis/isis_en.htm)
advanced (more specialized) spectroscopy software
 - Both by Christian Buil
- ImageTools by Peter Schlatter (private communication)
- SpectroTools by Peter Schlatter
<http://www.peterschlatter.ch/SpectroTools/>

Links

- SonotaCo Forum
<http://sonotaco.jp/forum/viewtopic.php?t=3065>
- Thorlabs grating
http://www.thorlabs.de/newgrouppage9.cfm?objectgroup_id=1123
- FMA Spectroscopy page
<http://www.meteorastronomie.ch/spektroskopie.html>
- Italian Meteor Network: <http://meteore.uai.it/>
- All documentation on calibration: <https://meteorspectroscopy.org>
(personal website, with papers and presentation from IMC 2016):

Acknowledgment

- FMA (division of Swiss (Amateur) Astronomical Society) for data, discussion
 - Jonas Schenker, Roger Spinner (website, database)
 - Network of stations (Photo, Video, All sky fireball detection, Radio, Seismic), complementing Spectroscopy
 - Linked with EDMOND database
- Peter Schlatter (Image tools)
- Koji Maeda (HD color videos)

Thank you!