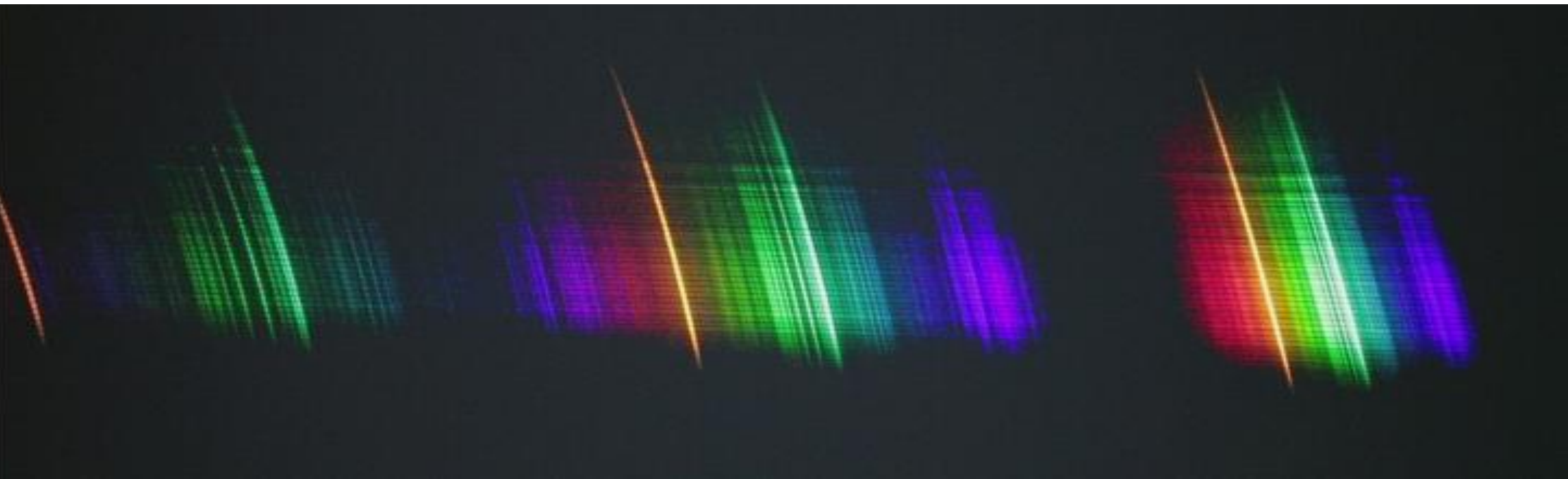


Meteor Spectroscopy, 2016

Martin Dubs, images by Koji Maeda

SAG, FMA

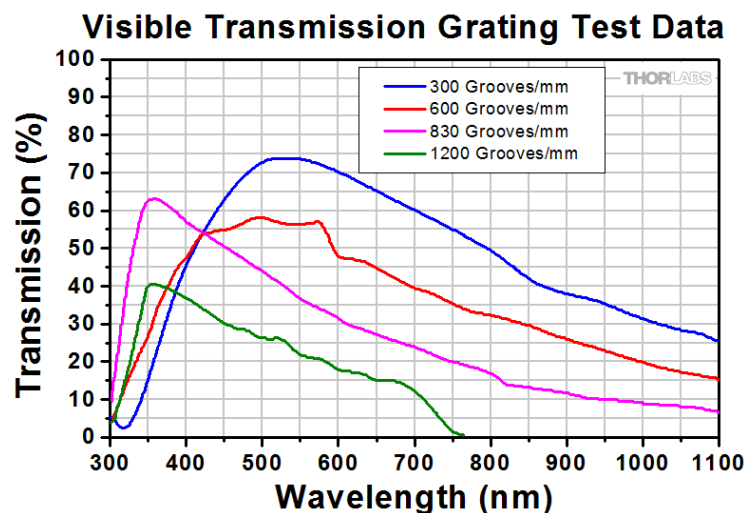


Inhalt

- Meteorspectroscopy
 - Hard- Software
 - Wavelength calibration
 - Spectrum extraction
 - Instrument response
 - Summary

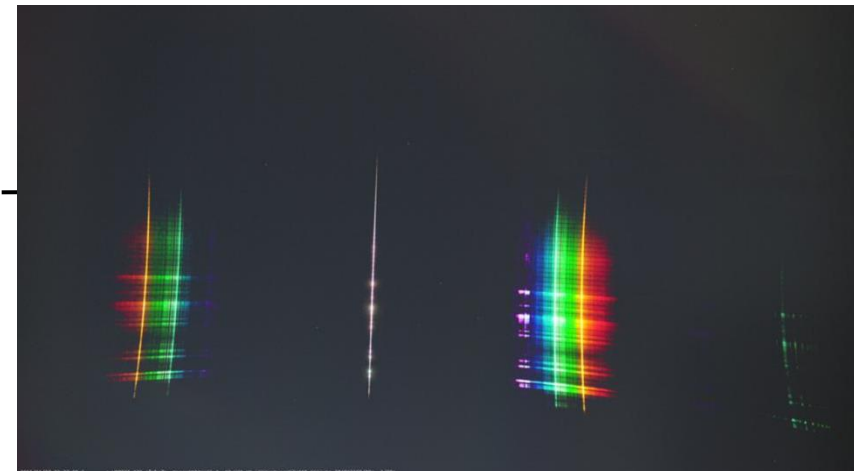
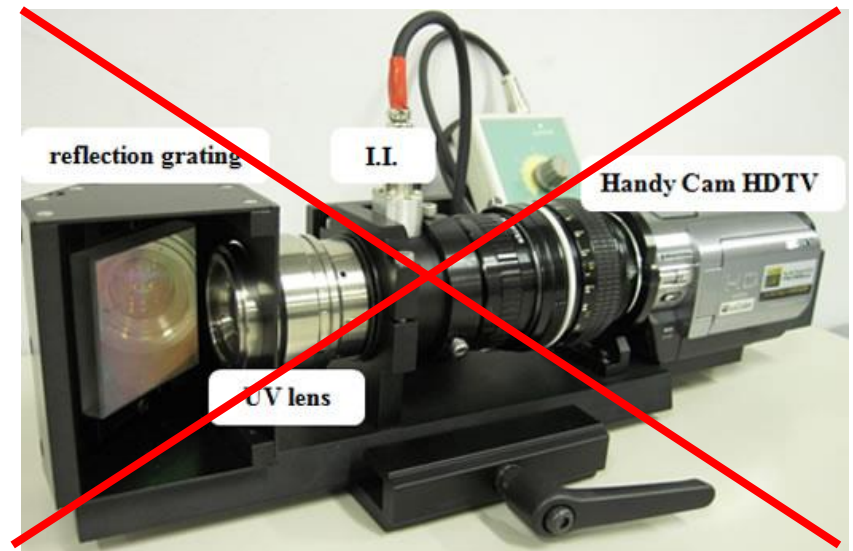
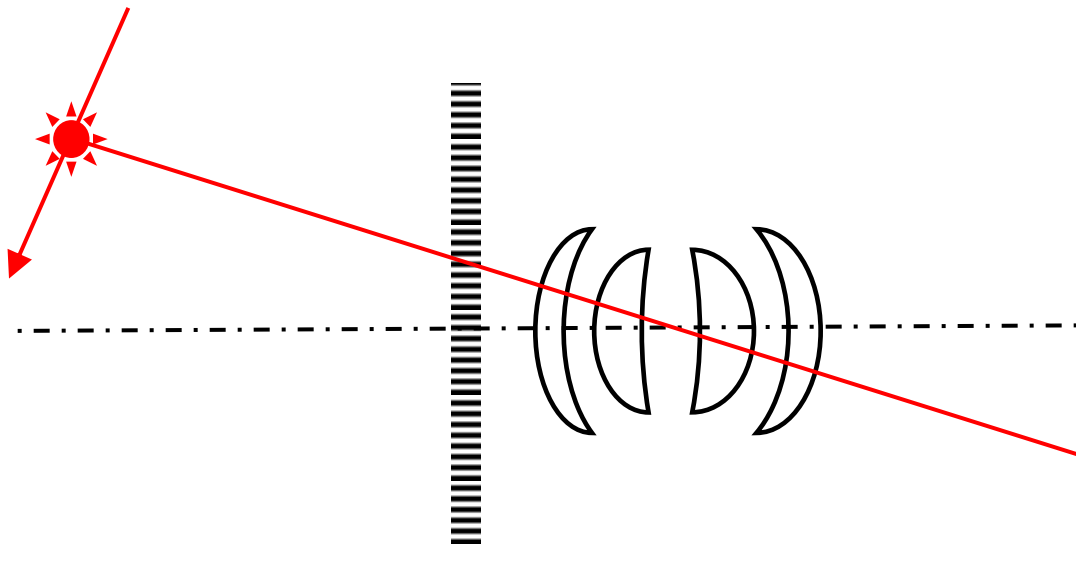
Hardware

- Watec 902H2 ult. Computar HG2610AFCS-HSP F/1 2.6mm fl
- 902H2 ultimate (spectroscopy) Tamron 12VG412ASIR F/1.2, $\approx 7\text{mm}$ fl
- 2nd camera with transmission grating for spectroscopy
Thorlabs
300 L/mm $\rightarrow$ 600 L/mm



Grating in front of lens

- 300 – 600 L/mm transmission grating
- Grating perpendicular to optical axis!



Vector theory wavelength calibration

- Grating in front of lens perpendicular to optical (z-)axis

- Unit vector (A B C) for incident direction

- Diffracted beam

$$A' = A + m\lambda G \quad (x\text{-axis})$$

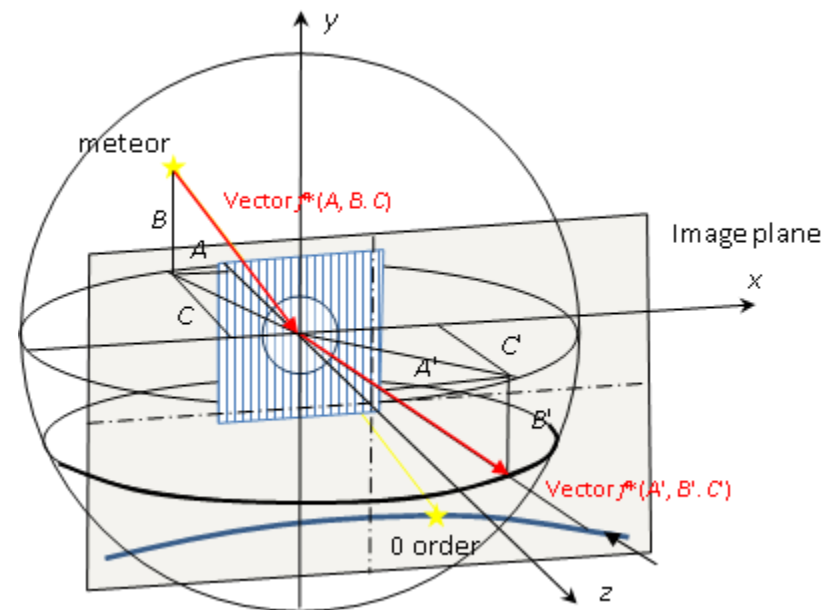
$$B' = B \quad (y\text{-axis})$$

$$C' = \sqrt{1 - A'^2 - B'^2}$$

- Spectrum on CCD plane

- Nonlinear dispersion
- Hyperbolic curvature

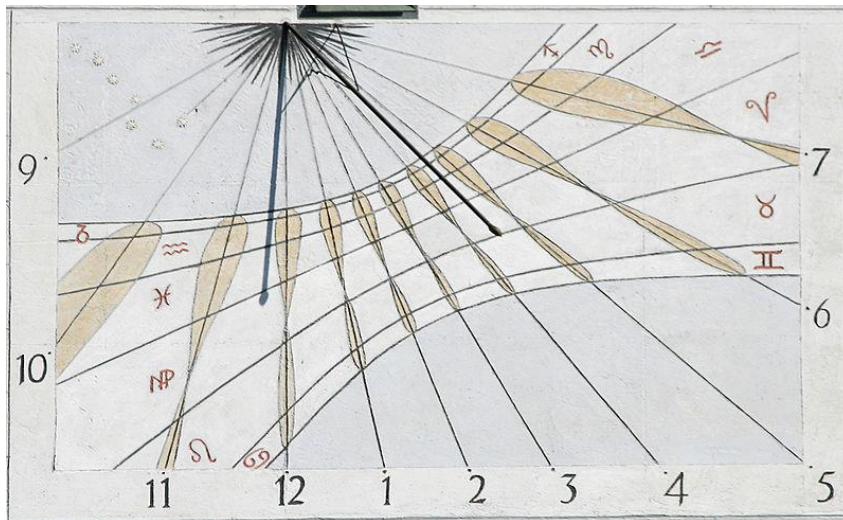
- Spectrum straight linear in A', B'



Gnomonic and orthographic projection

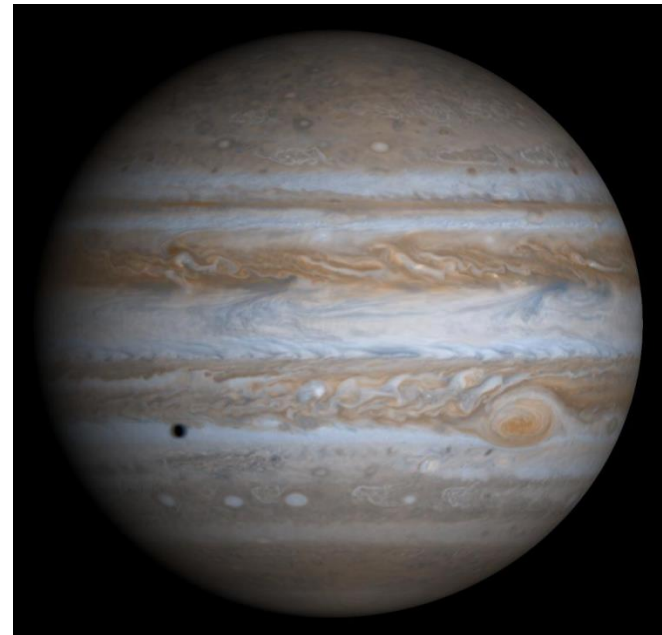
■ Gnomonic, TAN

- $R = f \cdot \tan(\rho)$
- Great circles $\rightarrow$ straight
- Optimum for path, radiant
- Latitude circles $\rightarrow$ hyperbola



■ Orthographic, SIN

- $R = f \cdot \sin(\rho)$
- Great circles $\rightarrow$ ellipses
- Latitude circles $\rightarrow$ straight
- Optimum for spectroscopy



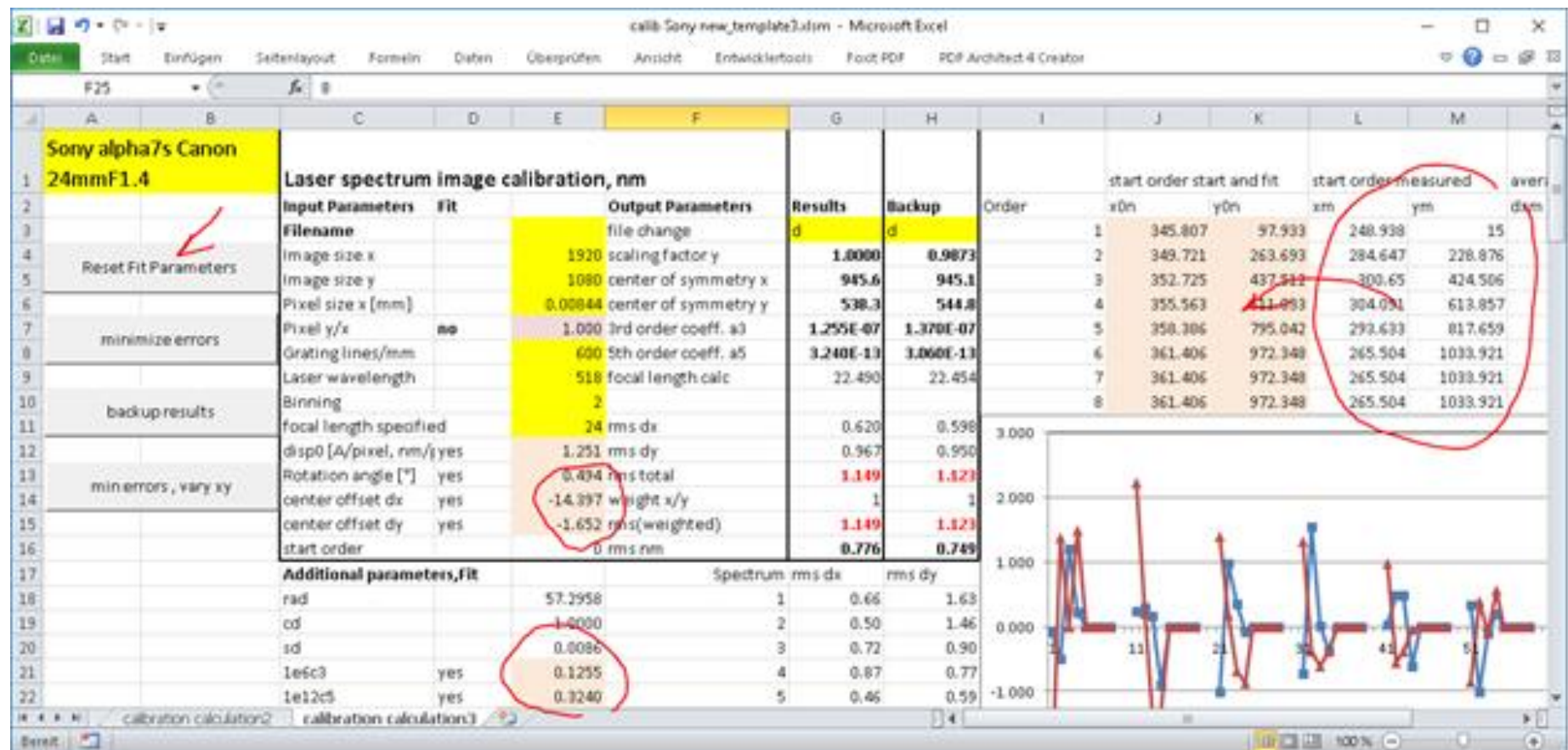
Calibration image

- Watec 902 H2 ultimate
- Tamron 12VG412ASIR,
- grating 600 L/mm
- Violet laser 405 nm

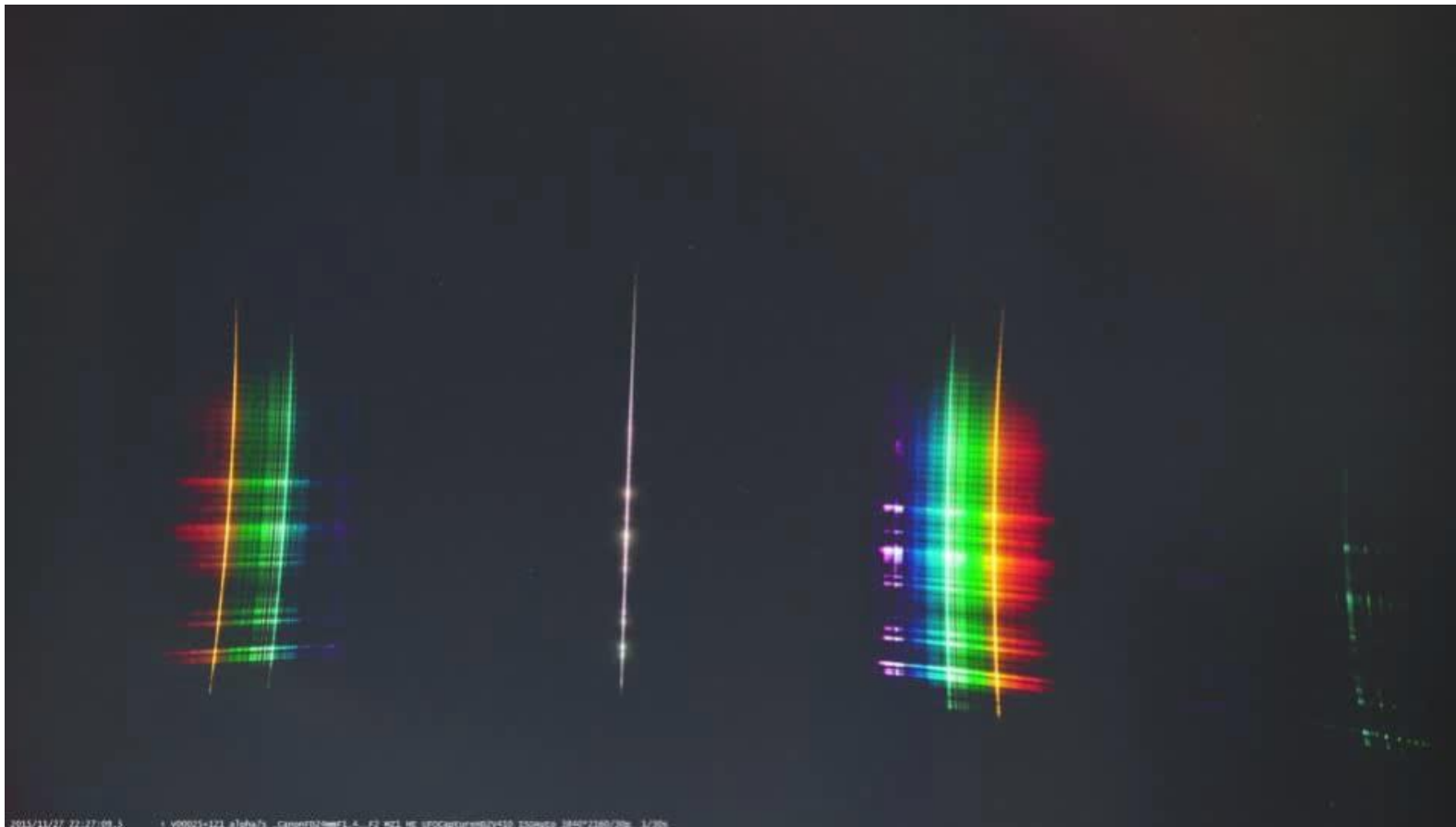


Calibration

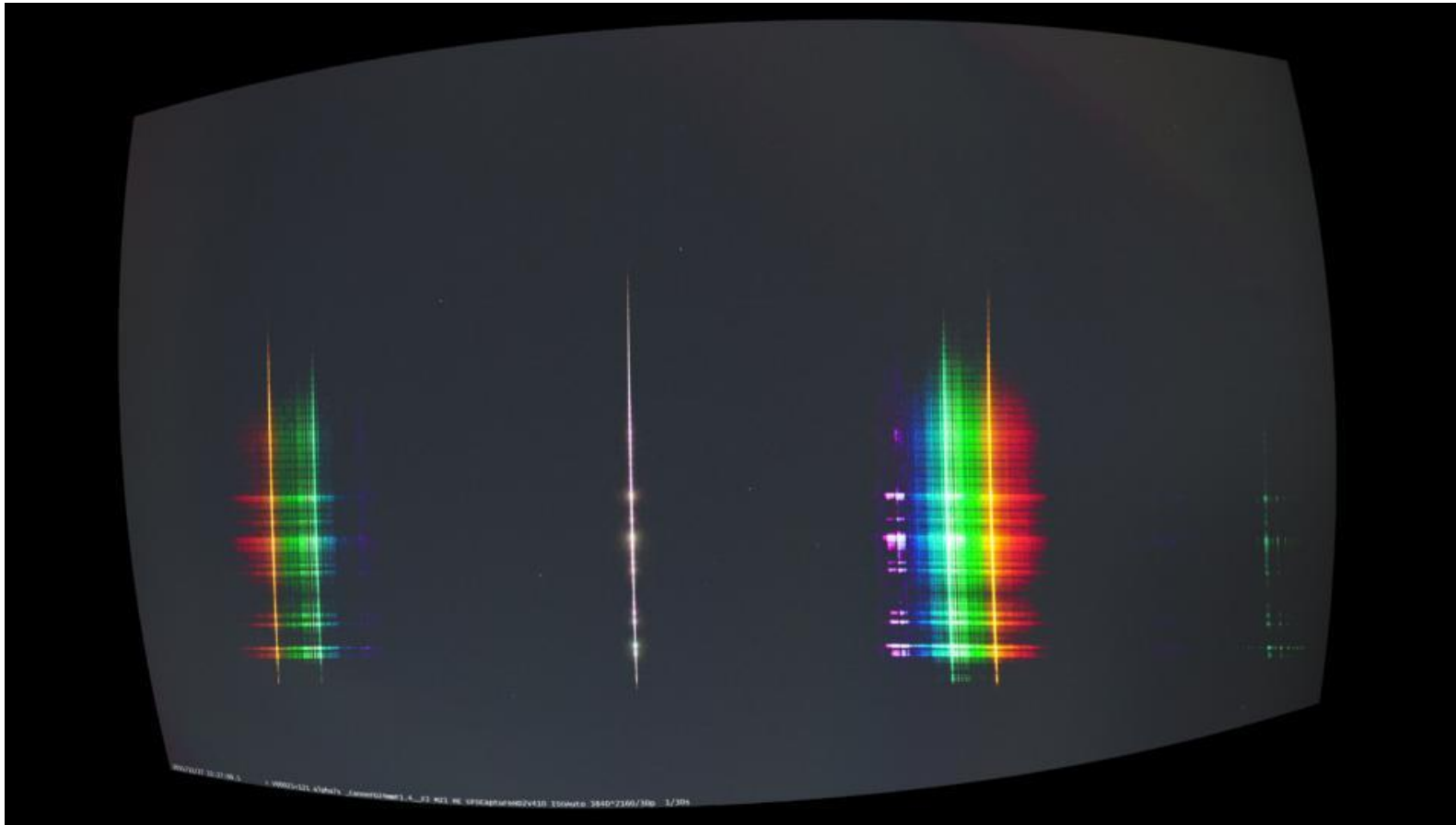
■ EXCEL worksheet



Orthographic transformation (1st price IMC)

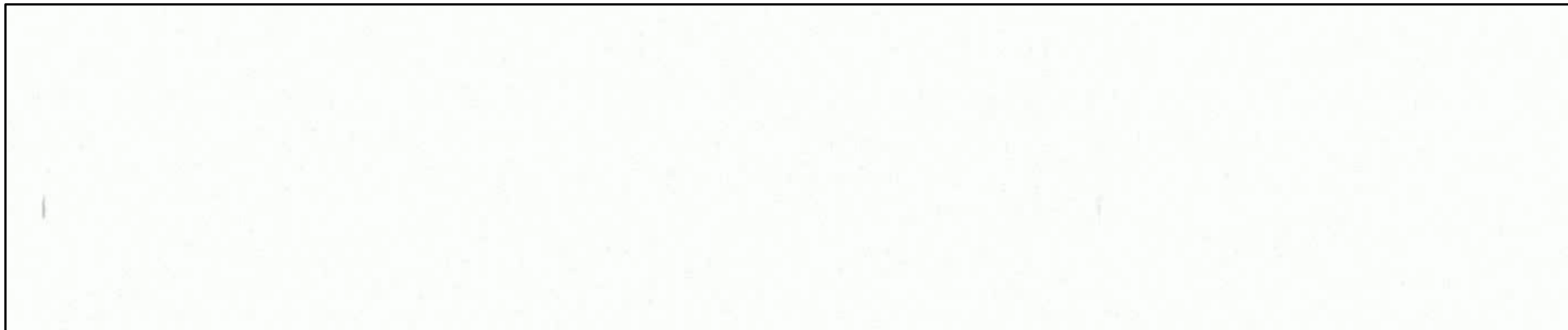


Orthographic transformation, result



Orthographic transformation, result

- Frames converted to b/w, linearized, registered, M20151127_222709



- color



Full processing

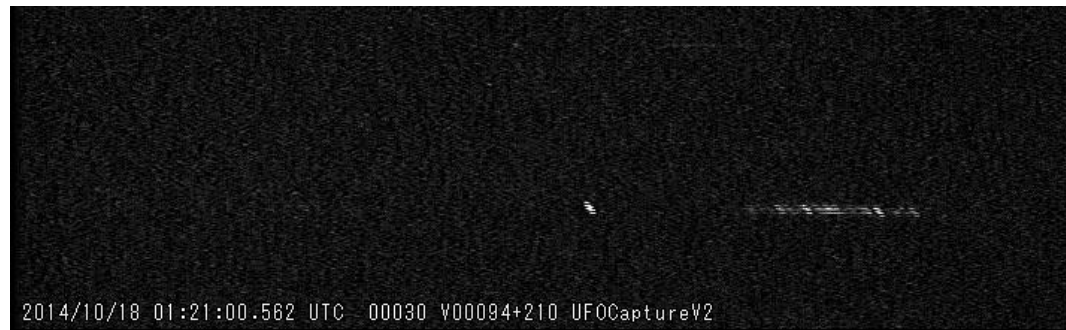
- Lens calibration ✓
- Preprocessing
 - Extract images from video
 - Background subtraction!
 - Vignetting, field of view
 - Correction for image transformation
- Apply image transformation
- Extract spectrum, calibrate wavelength
- Instrument response
 - Grating efficiency
 - Camera spectral sensitivity (lens, CCD)
 - Atmospheric transmittance

} flat field correction
in pre-processing

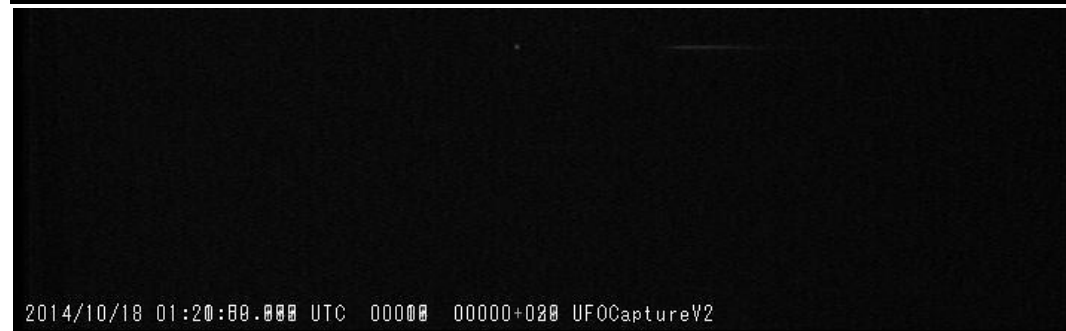
} instrument response

Preprocessing

- Extract image (i30)



- Background (ADD_MEAN)
better:
ADD2 MULT
< I1 ... I20 >

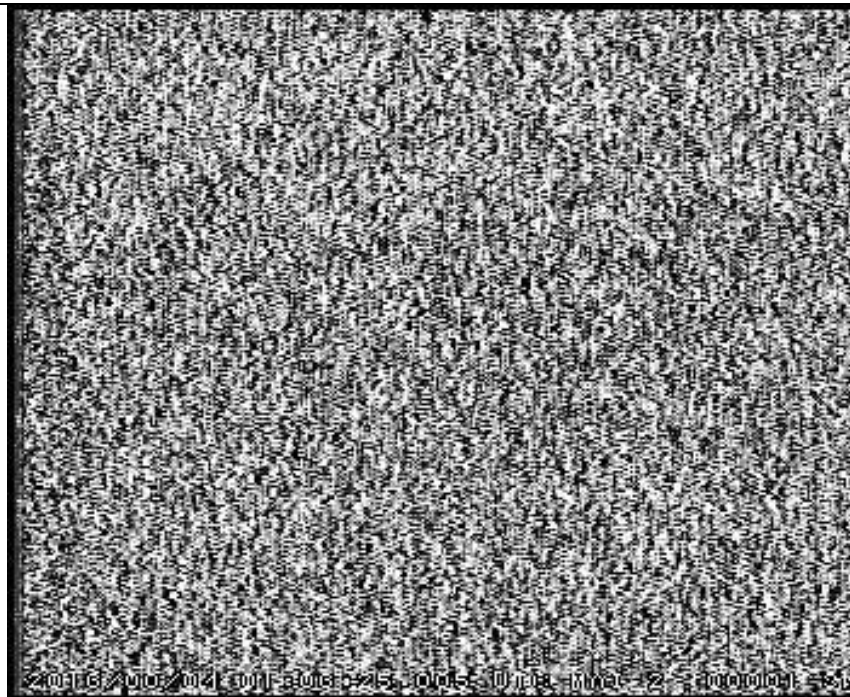


- Subtraction
SUB2

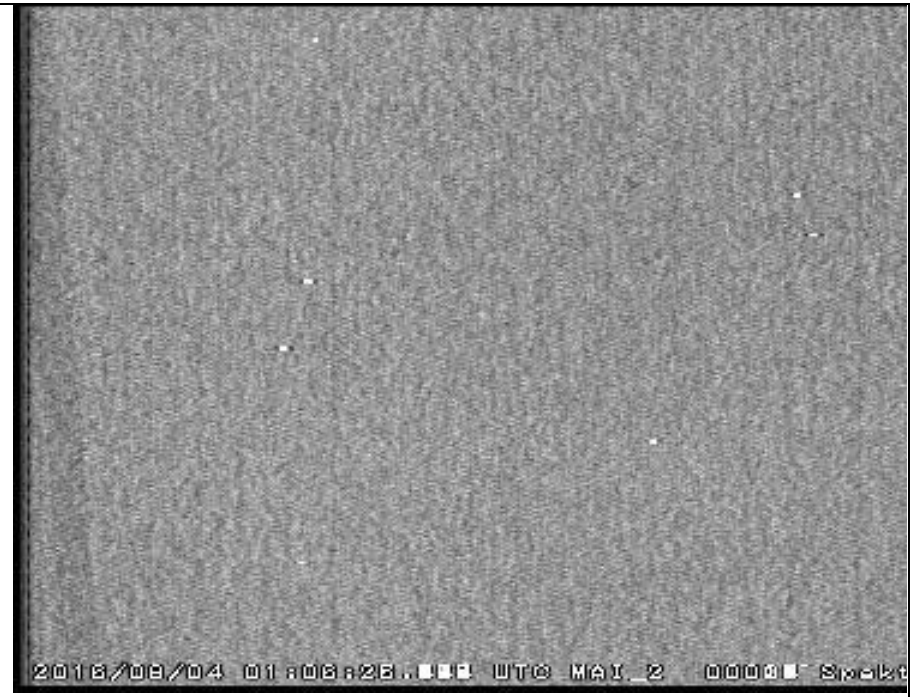


Preprocessing new

- RUN M_BACK: create background image
- RUN M_DARK: subtract background



M_1 visu 10 0 sigma = 4.4 ADU



Back visu 10 0 sigma = 0.75 ADU, some stars, hot pixels visible

Flat field

- Border not illuminated, same f-stop as meteor recordings (Nikon)

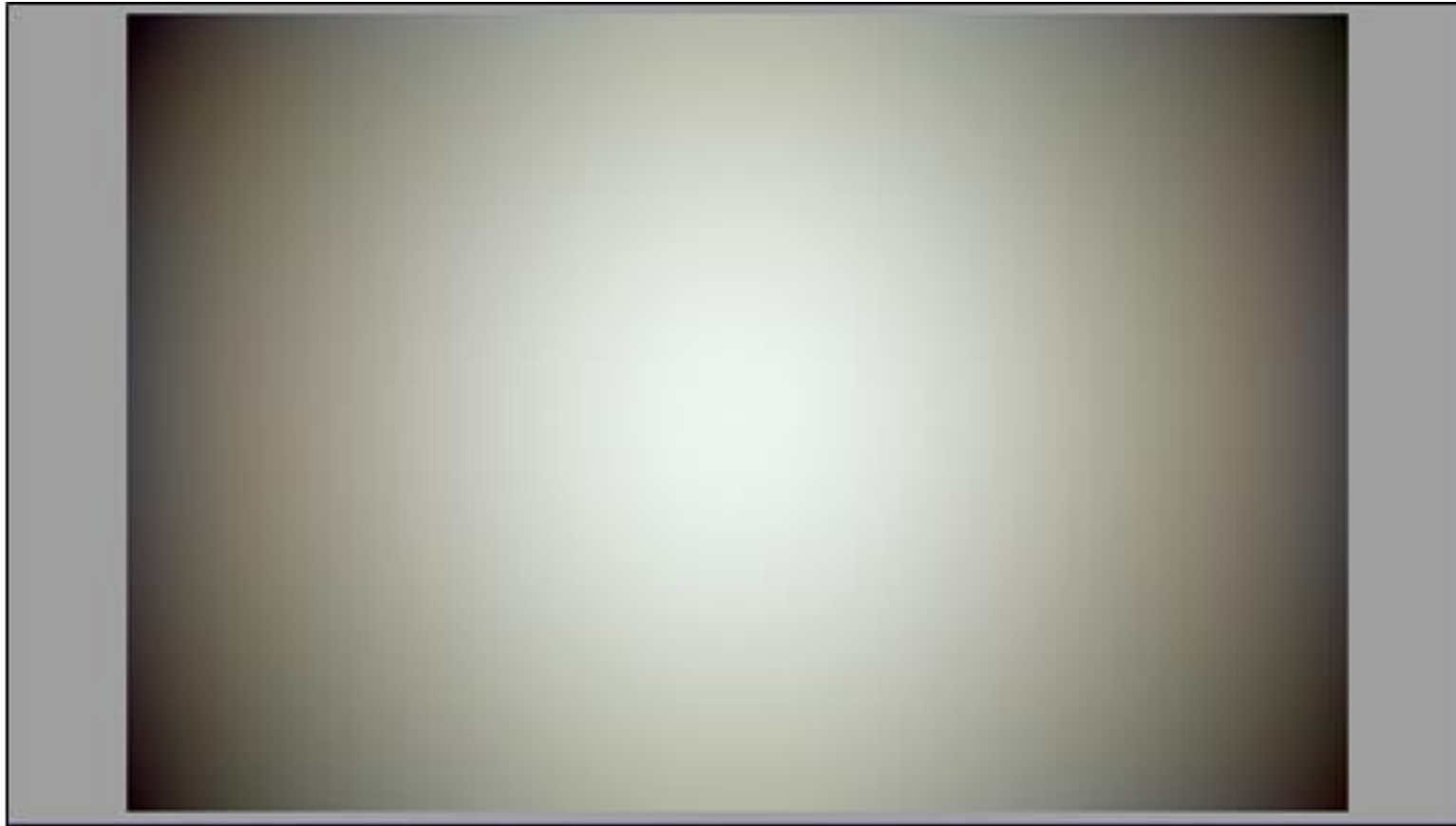
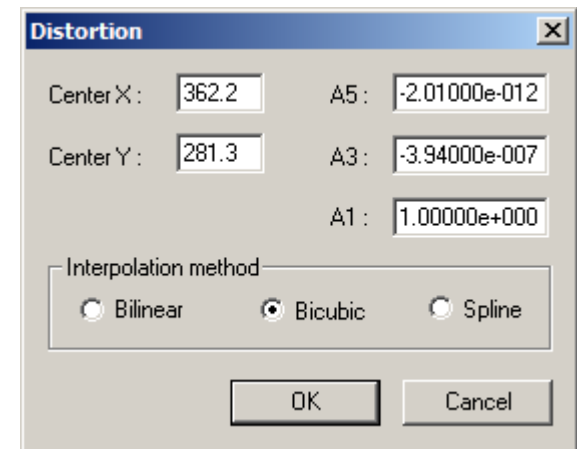
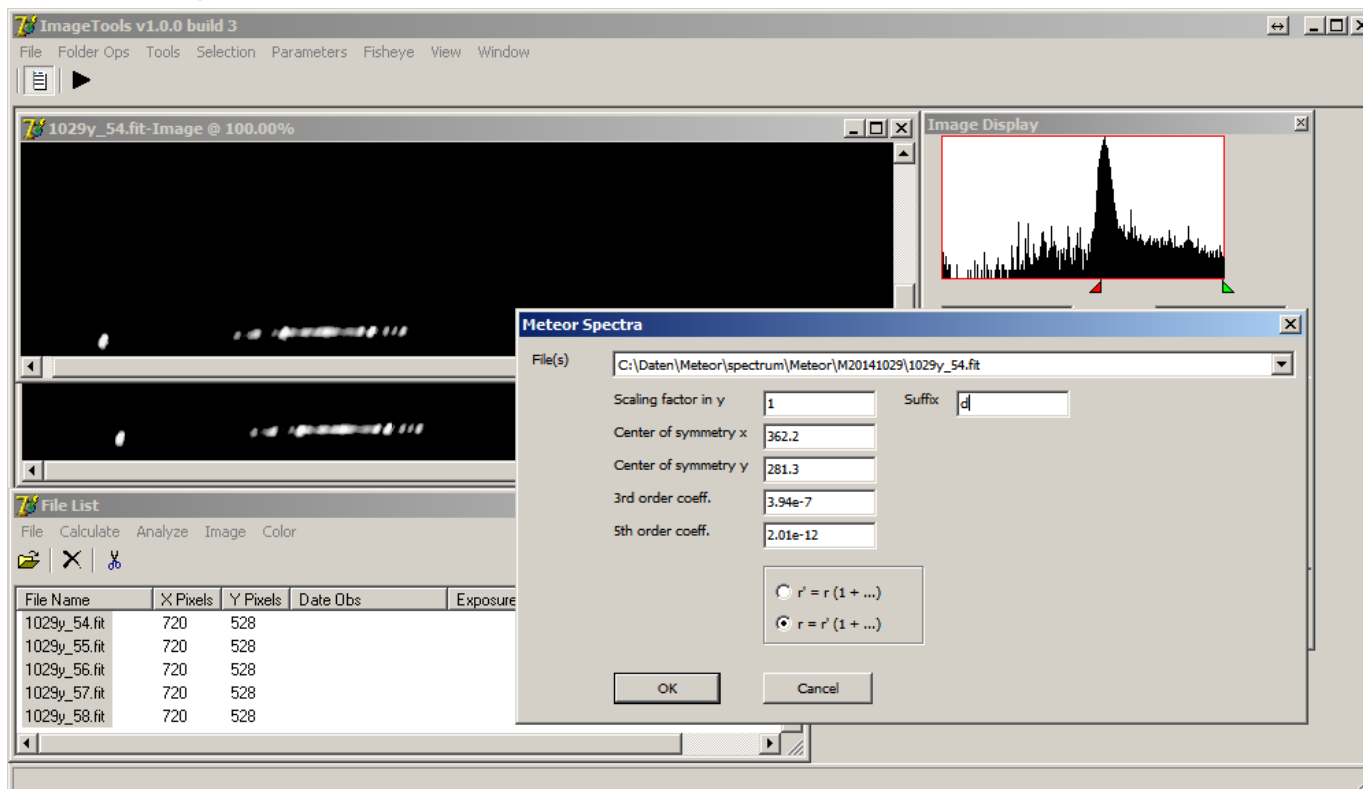


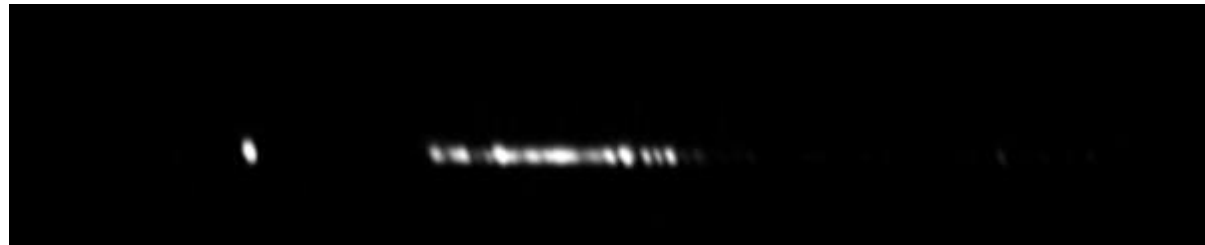
Image transformation

- Transformation to square pixels
- Transformation to orthographic projection IRIS → ImageTools →

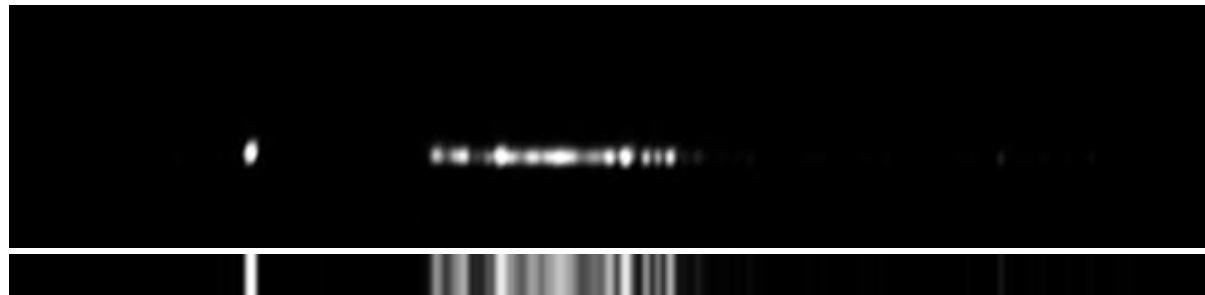


Processing 2, stacking

- REGISTER
ADD2
d1-15

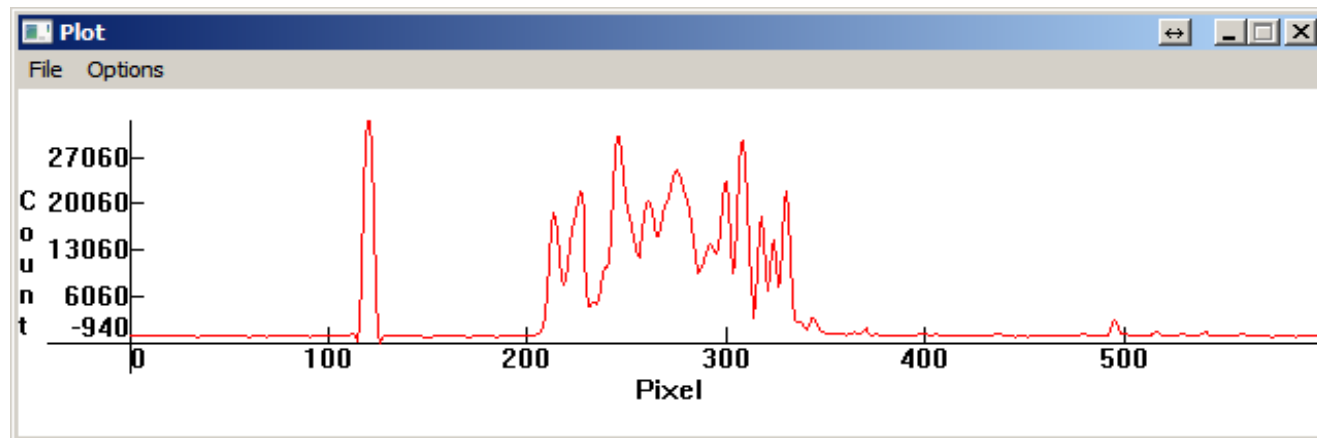


- SLANT 472 24



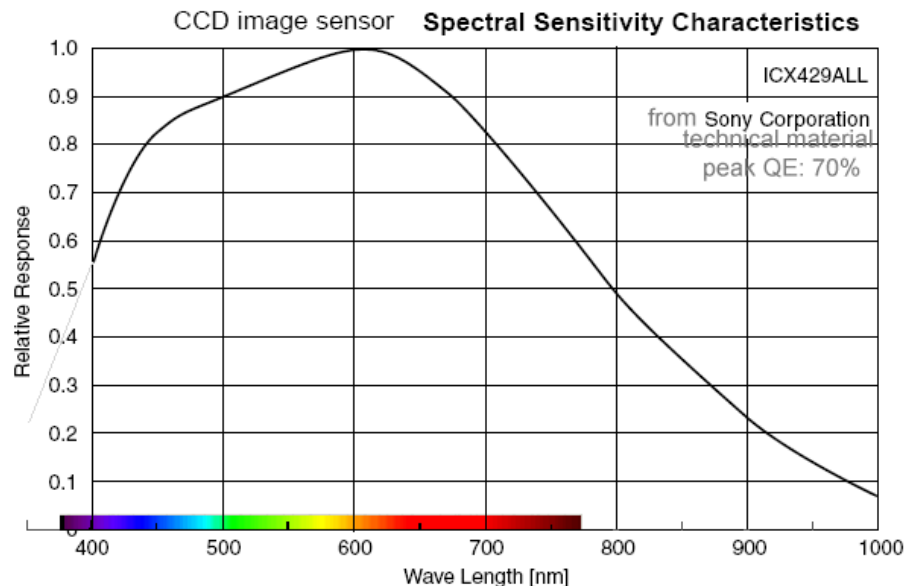
- L_ADD

- L_PLOT
save file.dat



Instrument response

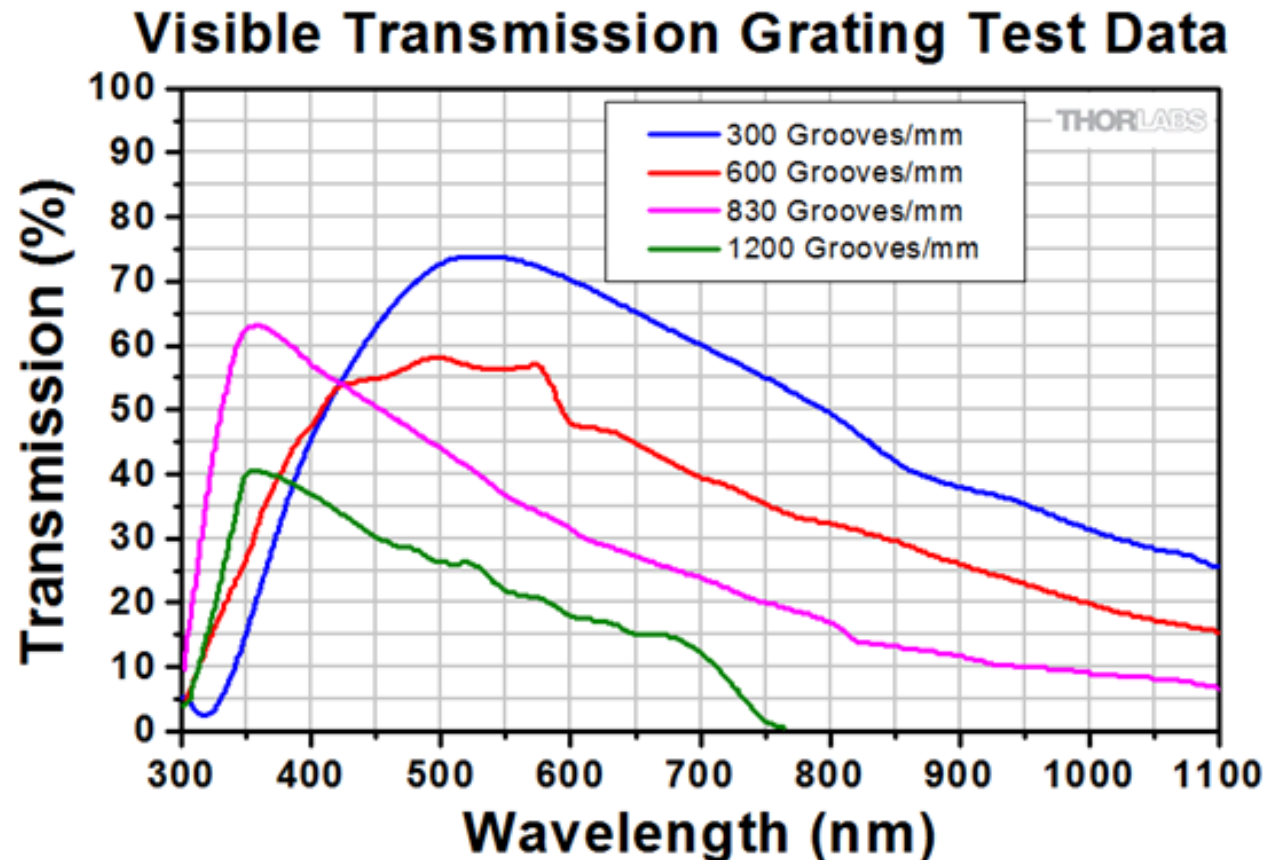
- Correction of
 - Camera spectral sensitivity (CCD, lens transmission)
 - Grating efficiency
 - Atmospheric transmission



ICX429ALL: Watec 902H2ultimate

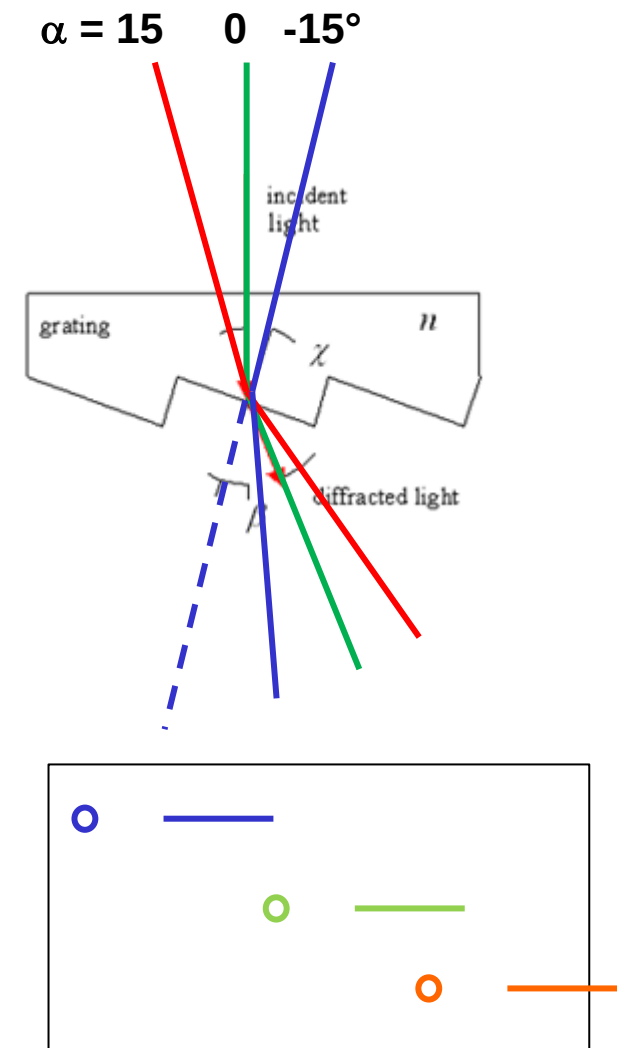
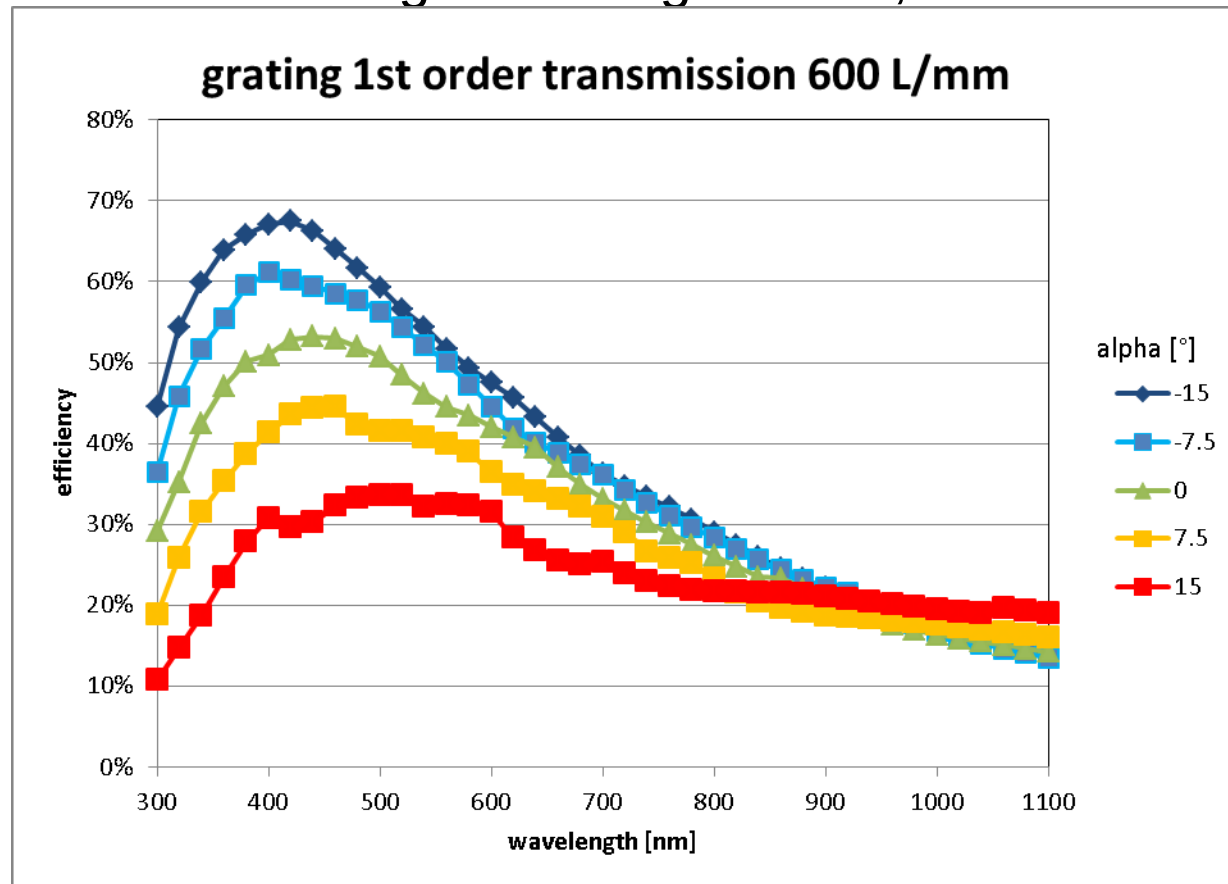
Grating efficiency

- Depends on grooves/mm and groove angle
- Thorlabs:



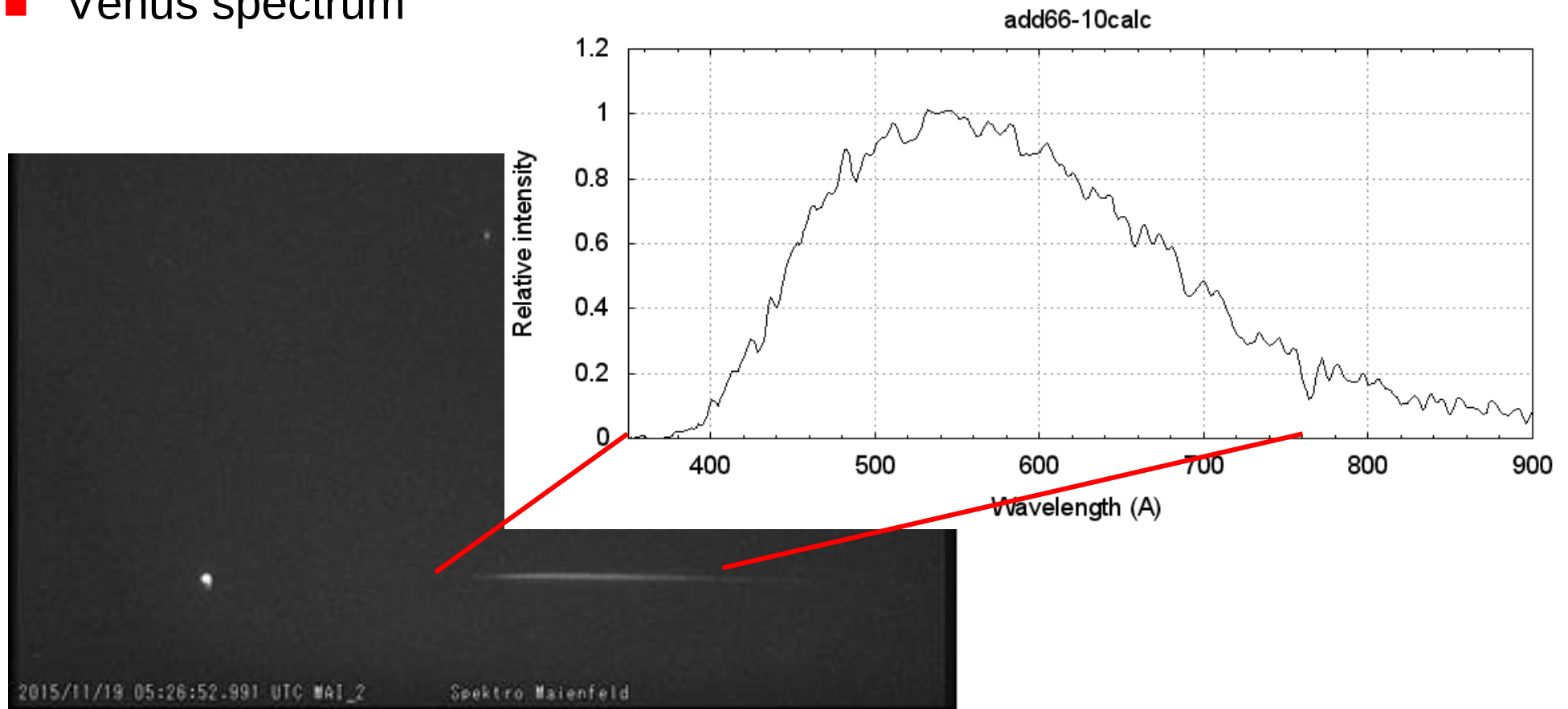
Grating efficiency 2

- Grating efficiency dependent on incident angle!
Calculated for groove angle 28.7° , $n = 1.52$



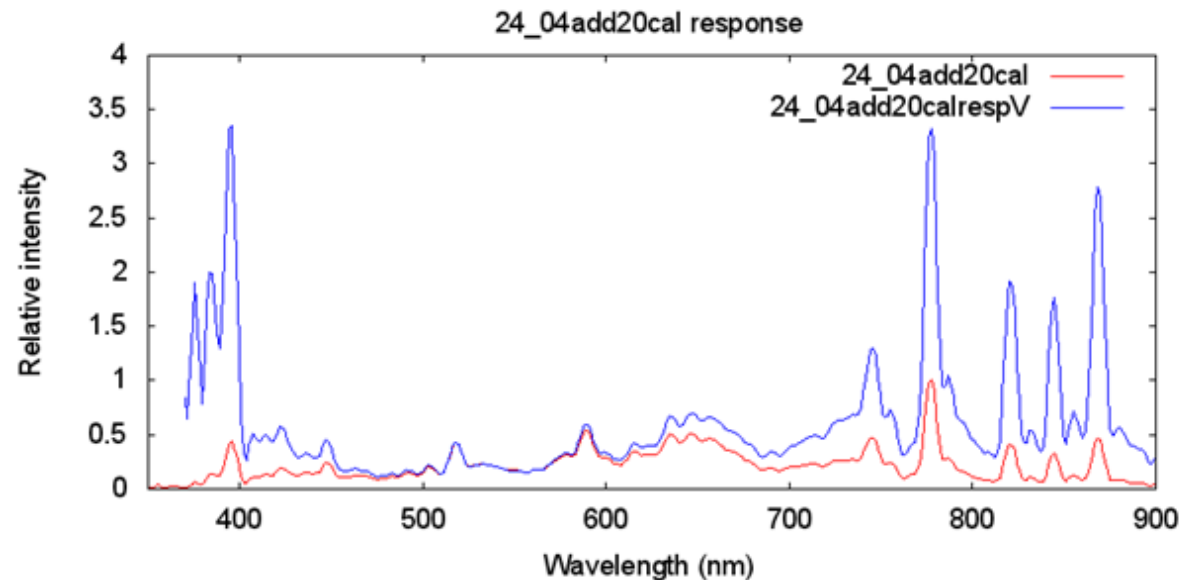
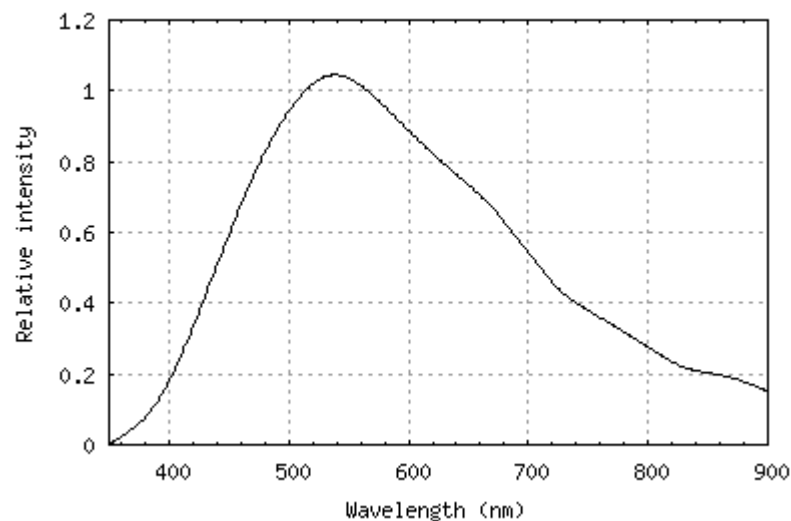
Measured spectrum, Watec 902H2ultimate

■ Venus spectrum



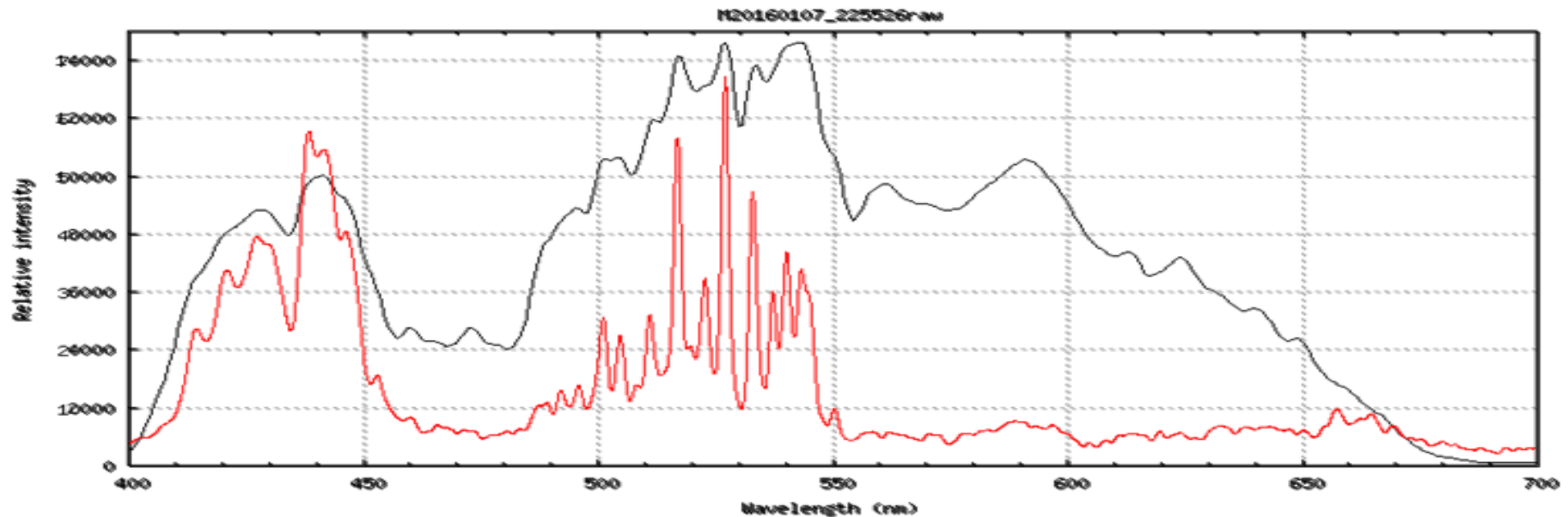
Instrument response

- Spectrum of known object (Venus, Sirius)
 - $IR(\lambda) = \text{measured spectrum}(\lambda) / \text{flux calibrated reference spectrum}(\lambda)$
- Meteor spectrum, wavelength calibrated $\rightarrow$ flux calibrated spectrum
 - $\text{Flux calibrated spectrum}(\lambda) = \text{meteor spectrum}(\lambda) / IR(\lambda)$



Saturation

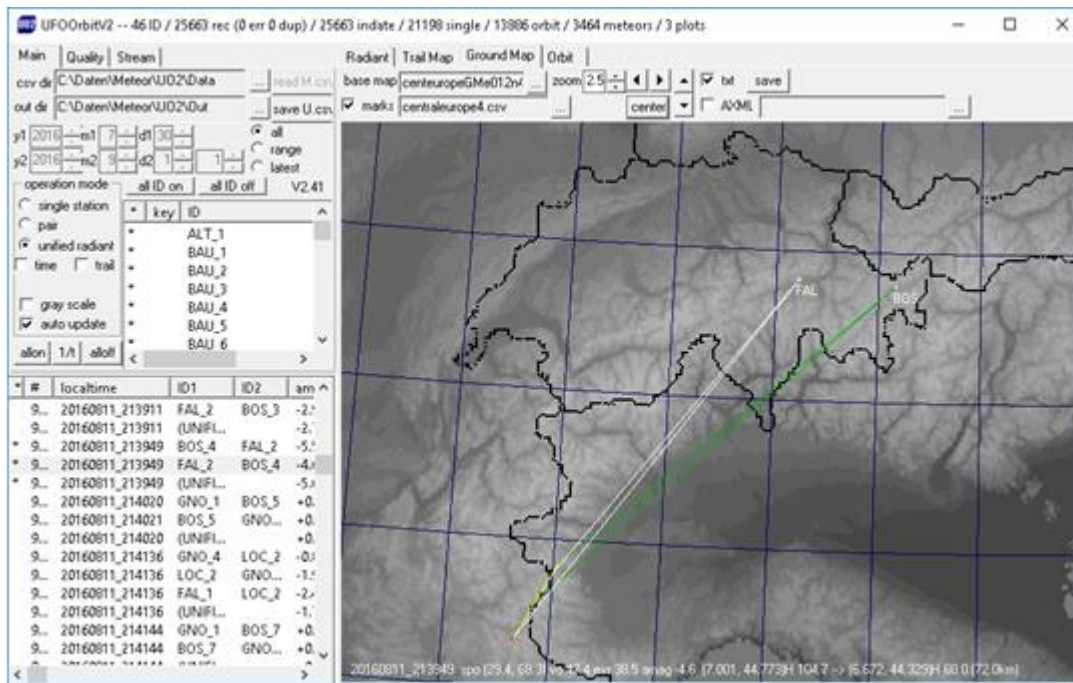
- 1st and 2nd order on same wavelength scale



Fachgruppe Meteorastronomie

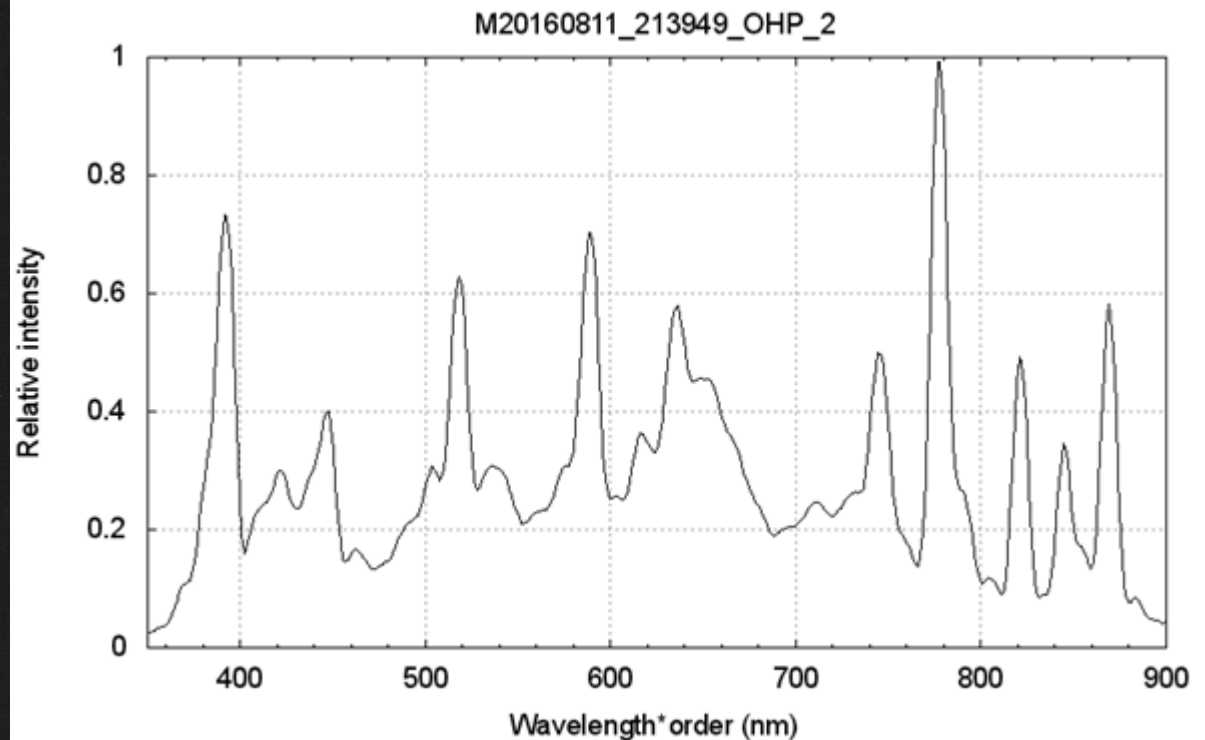
M20160811_213949_OHP_2, Per, $\cong -2.0m$

- Spectrum from OHP, triangulation from Switzerland



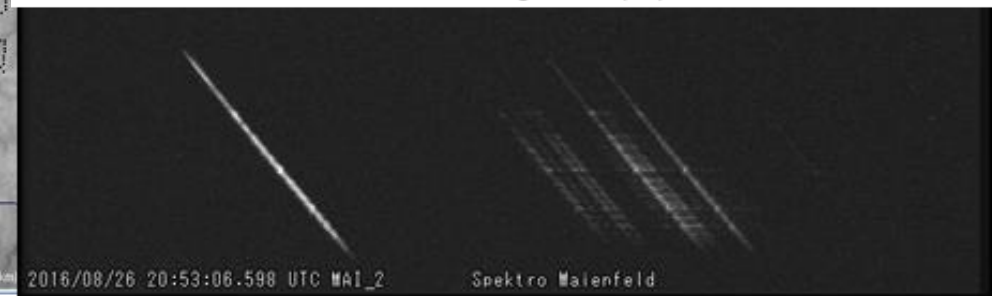
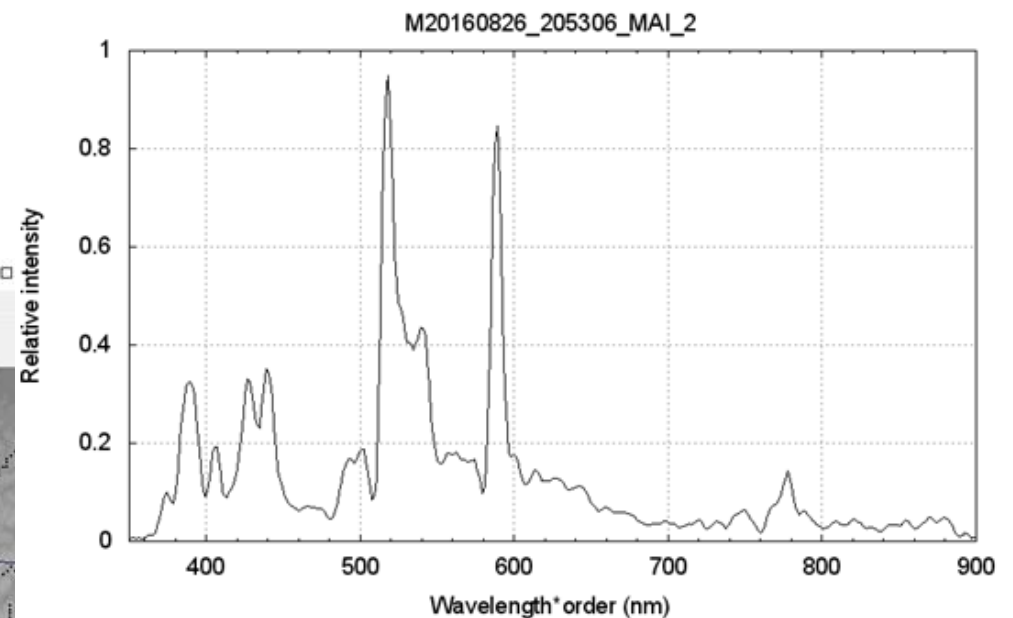
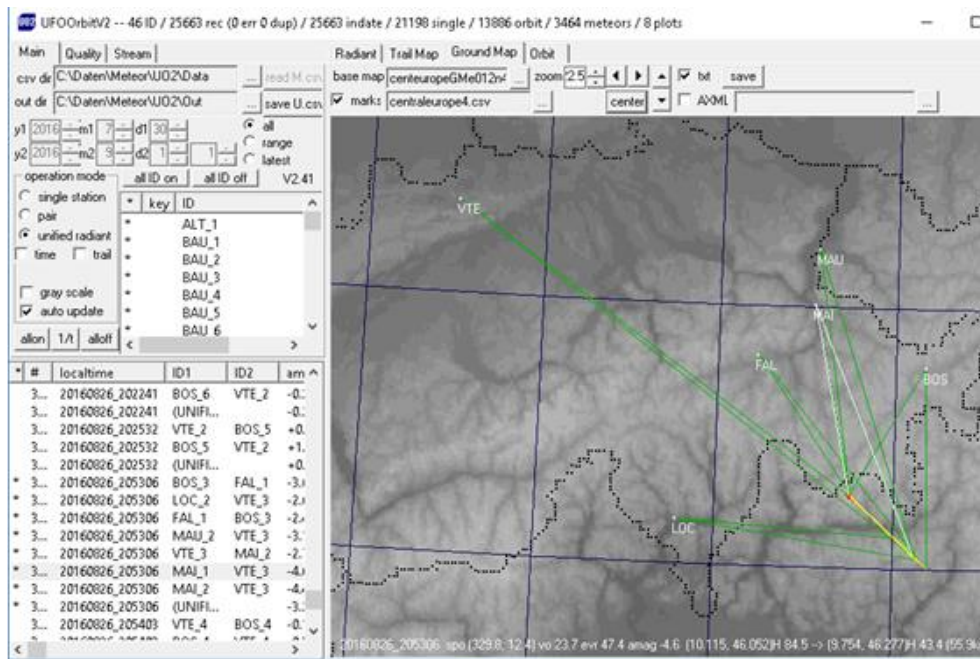
M20160811_213949_OHP_2, Per, $\cong -2.0\text{m}$

- Spectrum from OHP, triangulation from Switzerland



M20160826_205306_MAI_2, spo, -4.2m

- Very little atmospheric lines
- Observed by 6 stations



Trade-off focal length

- Short focal length
 - Large field of view
 - Low spectral resolution
 - Low sensitivity
 - Different orders visible
 - Many meteor spectra
- Grating with 300 to 1000L/mm mounted perpendicular to axis
- Long focal length
 - Small field of view
 - High spectral resolution
 - High sensitivity
 - Few meteor spectra of higher quality
- 6 – 12mm focal length for $\frac{1}{2}$ " CCD
- Limit magnitude -1 ... -2

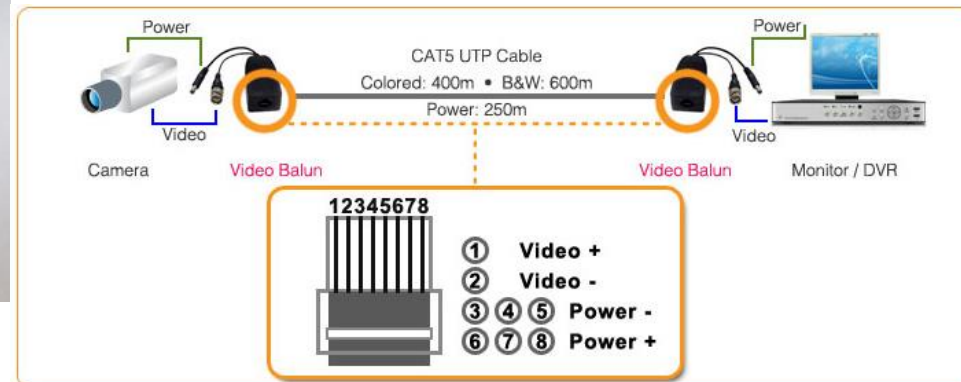
Waiting for next meteor spectrum, patience required!

Conclusion

- Grating mounted perpendicular to camera axis
- Image transformation gives linear, parallel spectra!
- Precise flux calibration depends on many factors, approximations used
- Looking for low cost, sensitive, high resolution, high dynamic range video camera
- Full format colour camera (e.g. Sony)
 - + Color → easy interpretation
 - + Orders can be separated
 - + High resolution
 - Bayer matrix lower sensitivity
 - Difficult to analyse (Instr. Resp.)
 - cost
- Video camera (e.g. Watec)
 - + High sensitivity
 - + Spectral range
 - + Low cost
 - Small field of view or
 - Low spectral resolution
 - Overlapping orders

Camera connection

- Video Balun 15\$/5 Paare (alibaba)
- Ethernet Flachbandkabel (Alibaba) (5\$/10m)
- USB 5 → 12V converter (ebay)



Spectrum recording and processing software

- UFO Capture for trigger and record video
(http://sonotaco.com/e_index.html)
- IRIS (<http://www.astrosurf.com/buil/us/iris/iris.htm>)
astronomical image processing and spectroscopy software
http://www.astrosurf.com/buil/iris/nav_pane/CommandsFrame.html
- ISIS (http://www.astrosurf.com/buil/isis/isis_en.htm)
advanced (more specialized) spectroscopy software
 - Both by Christian Buil
- ImageTools by Peter Schlatter (private communication)
- SpectroTools by Peter Schlatter
<http://www.peterschlatter.ch/SpectroTools/>

Acknowledgment

- FMA for data, discussion
 - Jonas Schenker, Roger Spinner (website, database)
 - Stefano Sposetti, Jose de Queiroz (equipment, data)
 - All others (data, discussions)
- Peter Schlatter (Image tools)
- Koji Maeda (HD color videos)

Links

- SonotaCo Forum
<http://sonotaco.jp/forum/viewtopic.php?t=3065>
- Thorlabs grating
http://www.thorlabs.de/newgrouppage9.cfm?objectgroup_id=1123
- Linear wavelength calibration: <http://arxiv.org/abs/1509.07531> or
http://www.meteorastronomie.ch/images/Meteor_Spectroscopy_WGN4_3-4_2015.pdf
- IMC:
http://www.meteorastronomie.ch/images/20160605_Calibration_of_Meteor_Spectra_Dubs_IMC2016.pdf

Danke für Ihre Aufmerksamkeit!